

# Detailed Study of the Sediments Recording the Matuyama-Brunhes Geomagnetic Reversal

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## ABSTRACT

The purpose of the present study is to find the relationship of a geomagnetic field reversal to the changes in a foraminiferal fauna and the sedimentary environments. A fossiliferous marine sedimentary section of the middle part of the Kokumoto Formation continuously exposed at Kamiyanagawa in the Boso Peninsula, Chiba Prefecture, Central Japan, was chosen for the investigation of the behaviour of the geomagnetic field during its reversal between the Matuyama and Brunhes Polarity epochs which took place around 0.69 m.y. ago. The behaviour of the geomagnetic field was established as to the inclination, declination and intensity of the  $J_n/J_R$  recorded in the sedimentary rocks. Along the same section, analyses were made on the grain-size of the sediments, oxygen isotope ratio in the planktonic and benthonic foraminiferal tests, and the faunal assemblages of the planktonic and benthonic foraminifers.

The intensity of the magnetic field fluctuated with a period of approximately 6700 years and did not disappear during the reversal. The migration of the virtual magnetic north pole at the magnetic field reversal between the Matuyama and Brunhes Polarity epochs occurred roughly along the meridian of 120° E long. in the period of 4700 years. In the same duration the depth of the sea water, under which the sediments accumulated, fluctuated within the range of 200-300 m with a period of approximately 7000 years. No particular relation between the magnetic field reversal and fluctuation of the water depth was recognized. The result of paleotemperature and faunal analyses of the foraminifers showed that the magnetic field reversal at the Matuyama Brunhes Polarity Epoch boundary may have been related to the changes in the circulation pattern of the upper watermasses which caused more influx of cold water into this area in the northwestern Pacific as compared with the period before the geomagnetic polarity change.

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## INTRODUCTION

Geomagnetic reversals during the last 5 million years have been recognized by magnetic measurements of volcanic rocks of known radiometric age (Cox *et al.*, 1963, 1967; Cox, 1968, 1969). The standard geomagnetic time scale has been extended to the records in sedimentary sequences from deep-sea cores (Opdyke *et al.*, 1966; Hays *et al.*, 1969), as well as those exposed on land (Nakagawa *et al.*, 1969, 1971; Niitsuma, 1970). The biostratigraphic studies of microfossils in deep-sea cores with the geomagnetic time control have made some authors to suggest that relation between the geomagnetic field reversals and the observed faunal appearances and extinctions exists (Opdyke *et al.*, 1966; Hays and Opdyke, 1967; Hays *et al.*, 1969). Others, however, (Waddington, 1967; Black, 1967; Harrison, 1968), after quantitatively estimating the effects of direct and/or indirect radiations from cosmic rays, solar flares and solar wind, have concluded that they can not be significant factors causing the appearance or extinction of marine species have if their amount increases during the reversals.

The appearance and extinction of some marine species may be caused by such biological changes that happen in an ecosystem such as changes in a part of the food chain, in

the ecologic position of the species, etc. They are intimately related to changes in the physical and chemical environments including the climate and circulatory systems in the ocean.

The present study aims to examine the relationships between geomagnetic reversals and changes in sedimentary environments with special reference to Foraminifera and to determine if any causal, whether relationships direct or indirect exist. A continuously exposed fossiliferous marine section has been chosen to see the behaviour of the magnetic field reversal between the Matuyama and Brunhes Polarity epochs that took place around 0.69 m.y. ago. Since the rate of sedimentation of the neritic sequence is much higher than those in the deep-sea, the geomagnetic records in detail the behavior of the magnetic field during the reversal. The rate of sedimentation has been estimated for this section using geomagnetic chronology (Nakagawa *et al.*, 1969), so that it is possible to estimate the amount of time which elapsed during the reversal and other related changes. Along with the geomagnetic measurements, the following analyses have been made. They are a sedimentological analysis of the sediments, paleotemperature measurements of planktonic and benthonic foraminiferal tests by means of oxygen isotopic ratio and faunal analyses of the benthonic and planktonic foraminiferal fauna.

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#### NOTE ON THE GEOLOGICAL HISTORY AND THE LATE CENOZOIC GEOMAGNETIC STRATIGRAPHY OF THE BOSO PENINSULA

The Boso Peninsula is situated in the southeastern part of the Kanto Region, central Japan (Fig. 1).

The stratigraphy of the Neogene and Pleistocene rocks exposed in the peninsula is summarized in Table 1, following the scheme by Nakagawa (1962).

The total thickness of the Nakahara and younger formations exceeds 6000 meters. They are all of marine origin. The series is continuous in sequence except for the sedimentary break between the Anno and Kurotaki formations; the hiatus is estimated to be equivalent to about 150 m thick sediments of the surrounding area. The marine Neogene and Pleistocene formations of the Boso Peninsula consist largely of alternations of siltstone and sandstone. They frequently intercalate pyroclastic layers by which detailed tephro-

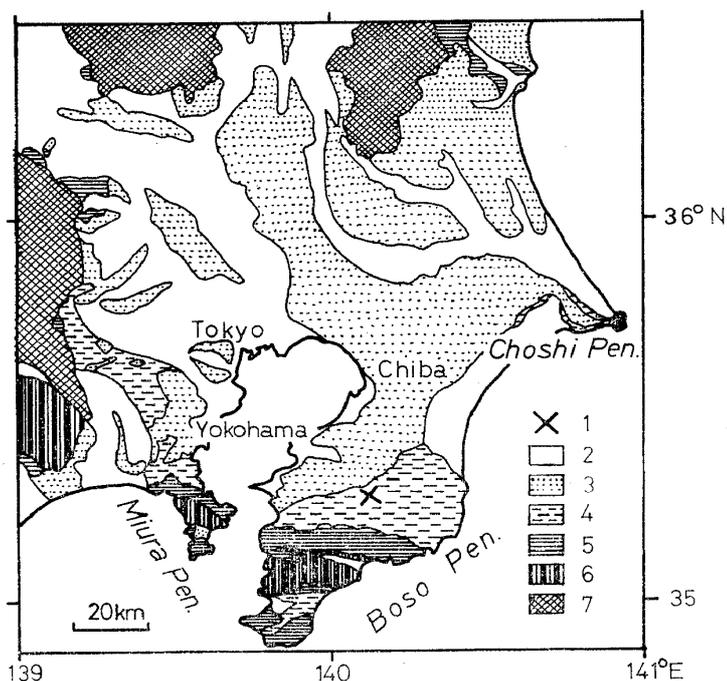


Fig. 1 Geologic sketch map of Kanto Region, Central Japan.

- |                            |                              |                 |
|----------------------------|------------------------------|-----------------|
| 1: Sampling position       | 2: Younger sediments         | 3: Narita Group |
| 4: Akimoto and Seki groups | 5: Toyooka and Sakuma groups |                 |
| 6: Hota and Mineoka groups | 7: Pre-Tertiary rocks        |                 |

Table 1 Stratigraphy of Neogene and Pleistocene rocks exposed in the Boso Peninsula.

	Kioroshi Formation
	Kamiwahashi F.
	Semata F.
	Yabu F.
Narita Group	Jizodo F.
	Kasamori F.
	Mandano F.
	Chonan F.
	—————(conformity)—————
Akimoto G.	Kakinokidai F.
	Kokumoto F.
	Umegase F.
	—————(conformity)—————
Seki G.	Otadai F.
	Kiwada F.
	Kurotaki F.
	—————(unconformity)—————
Toyooka G.	Anno F.
	Kiyosumi F.
	—————(conformity)—————
Sakuma G.	Amatsu F.
	Nakahara F.
	—————(unconformity)—————
Hota G.	
	—————(fault)—————
Mineoka G.	

chronology of the sedimentary rocks has been established (Mitsunashi *et al.*, 1959, 1961). The marine formations are fossiliferous and a number of paleontological and biostratigraphical studies have been made dealing with molluscs (e.g. Hatai, 1958; Ogose, 1961) and

benthonic and planktonic Foraminifera (e.g. Asano, 1938; Saito, 1963; Aoki, 1968).

The Mineoka and Hota groups are distributed along the axis of the anticlinal zone of the southern part of the Boso Peninsula (Fig. 1), and are intensely faulted and folded and intruded by basic and ultrabasic igneous rocks. The Nakahara Formation, unconformably overlying the Hota Group, is also folded and faulted in the anticlinal zone. The Amatsu Formation directly covers the Hota Group and together with the overlying formation is gently folded. The Nakahara Formation includes a horizon correlated with the *Globigerinatella insueta*/*Globigerinoides bisphericus* Subzone (Saito, 1963), and yielded larger Foraminifera as *Amphistegina*, *Lepidocyclina*, *Miogyopsina*, as well as Mollusca and other fossils (Hanzawa, 1931). The Kurotaki and younger formations dip homoclinally north-westwards. The Kioroshi Formation, the youngest of the Narita Group, is subhorizontal and is distributed not only in the peninsula but also widely in the central part of the Kanto Region. The Kioroshi Formation is conformably overlain by the clay of volcanic ash origin of the Shimosueyoshi Volcanic Ash covering the Shimosueyoshi Terrace formed during the last high sea-level and warm water temperature stage in the Late Pleistocene (Nakagawa, 1965).

In the later phase of the accumulation of the Toyooka Group, a zone with axis trending in WNW-ESE direction had been anticlinarily uplifted in the preceding marine sedimentary basin of the southern part of the Kanto Region, in which the Toyooka Group was deposited. In the Kazusa Sedimentary Basin developed succeedingly on the north of this anticlinal zone, thick marine Pliocene and Pleistocene sediments of the Seki and Akimoto groups were accumulated. The marine sedimentation appears to have been continuous in the central area even during the later transformation of the trough-shaped Kazusa Sedimentary Basin into the doubly plunging synclines called the Kanto Tectonic Basin, in which the Narita Group was deposited.

An attempt was once made by Kawai (1951) to measure the Natural Remanent Magnetization (NRM) of the samples from the Late Cenozoic marine sediments of the Boso Peninsula. Recently, Nakagawa, Niitsuma and Hayasaka (1969) succeeded in measuring the Detrital Remanent Magnetization (DRM) of siltstones in the sequence from the Amatsu up to the Kioroshi formations in the peninsula (Fig. 2). The results correlate well with the scheme of geomagnetic field reversals given by Cox (1969).

Correlating with the geomagnetic scheme of Cox (1969), Nakagawa *et al.* (1969) were able to identify the boundary between the Brunhes and Matuyama Polarity epochs in the middle part of the Kokumoto Formation, that between the Matuyama and Gauss in the lower part of the Umegase Formation, and that between the Gauss and Gilbert in the lowermost part of the Otadai Formation.

Comparison of the paleomagnetic and microbiostratigraphic sequence between the Boso Peninsula and south Italy (Nakagawa, Niitsuma and Elmi, 1971) suggests that the time equivalent of the boundary between the Pliocene and the Calabrian in the Castella area is likely to be found in the upper part of the Umegase Formation.

The boundary between the Pliocene and Miocene is now interpreted to be the unconformity between the Anno and Kurotaki formations, if we accept the correlation between the planktonic Foraminifera zones and geomagnetic polarity sequence presented by Berggren (1969).

#### DEPOSITIONAL CONDITION AND RATE OF SEDIMENTATION OF SILTSTONE

Pliocene and Pleistocene marine formations in the Boso Peninsula consist mostly of alternations of sandstone and siltstone. Thin layers of tuff are intercalated in the

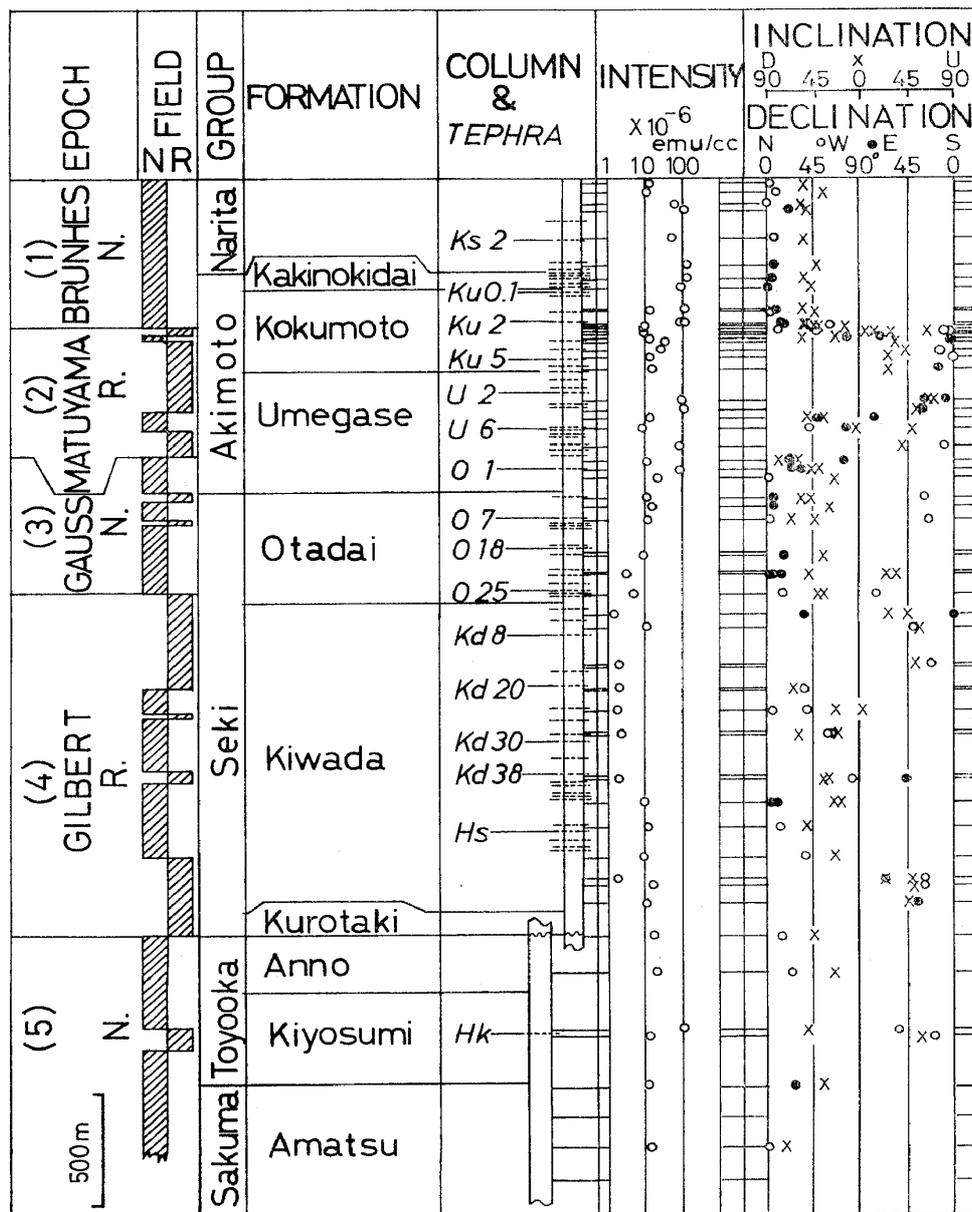


Fig. 2 Late Cenozoic geomagnetic stratigraphy in the central and northern parts of the Boso Peninsula.

The original record (Nakagawa *et al.*, 1969), particularly of those below the Gauss has been re-interpreted to divide them into two magnetic polarity epochs (Niitsuma, 1971) rather than four as was originally done.

siltstones and they are frequently useful as keys in tracing stratigraphic levels over widely separated sections. For example, a set of three 10–20 cm thick layers of tuff, named Ku 2 is intercalated in the middle part of a 150 m thick siltstone bed (Fig. 2) that forms a part of the alternation of siltstone and sandstone of the Kokumoto Formation, in the Yoro Valley of the peninsula. The Ku 2 can be traced for distance of 45 km from the west to east coasts in the central part of the Boso Peninsula (Mitsunashi *et al.*, 1959; 1961), its position relative to the Matuyama / Brunhes boundary has already been ascertained in the previous study (Nakagawa *et al.*, 1969).

The thickness of the siltstones interbedded with tuff layers, as well as the thickness of each tuff layer changes little in the western part of the Boso Peninsula, therefore it is inferred that the rate of sedimentation was mostly constant in this area. The depositional surfaces of the tuff layers are always flat and parallel with the bedding plane. These facts are taken as indication that the tuffs had been deposited on the horizontal sea bottom under a calm condition. Disturbance by burrowing organisms is sometimes observed along the upper surface of tuff layers, but the disturbance does not extend more than a few centimeters in depth, suggesting that the siltstones in which the tuffs are intercalated were deposited also on a horizontal plane at a constant rate of sedimentation under a calm condition. This observation holds true in this study (see Fig. 6).

Hirayama and Suzuki (1968) reported that there are changes in thickness of each members in the alternations of sandstone and siltstone of the Otadai Formation in an interval (3–5 m thick) between two key tuffs, O<sub>7</sub> and K<sub>3</sub>. On the basis of the observations made on this interval in the central part of the Boso Peninsula over the distance of 45 km, they found that the thickness of each siltstone bed changes a little; each sandstone bed, however, thickness as much as twice in the central part of the Boso Peninsula. This difference in the lateral changes of thickness between sandstone and siltstone seemed to suggest that the mechanisms of their deposition may have been different. The alternation beds in the Kokumoto Formation is of the same type as the Otadai Formation. In order to bear this possibility out, the sandstone layers alternating with siltstone were examined particularly along the magnetically studied section at Yanagawa (Fig. 5). The thickness measured in that section shows a logarithmic normal distribution, which suggest that the deposition of sandstone layers were governed by a single cause (Fig. 3). The sandstones were studied further as to their texture by means of the automatic grain-size analyser, made specifically by the present author (Niitsuma, 1971). The grain-size distribution patterns observed for sandstones thicker than 100 cm are constantly unimodal, with the maximal at  $2.6\phi$  (Fig. 4). The pattern coincide with what have been found for modern sands along the near shore zone of 10–13 m in depths both in Mutsu Bay and the northern part of the Pacific coast of Japan (Niitsuma and Mekata, 1971a, b). The grain-size dist-

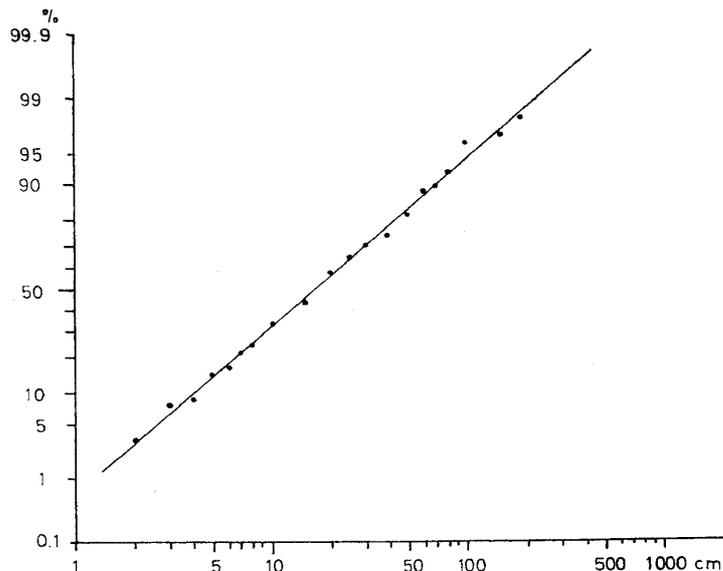


Fig. 3 Thickness in logarithmic scale of sandstones in the alternation of sandstone and siltstone of the Kokumoto Formation, plotted on probability paper.

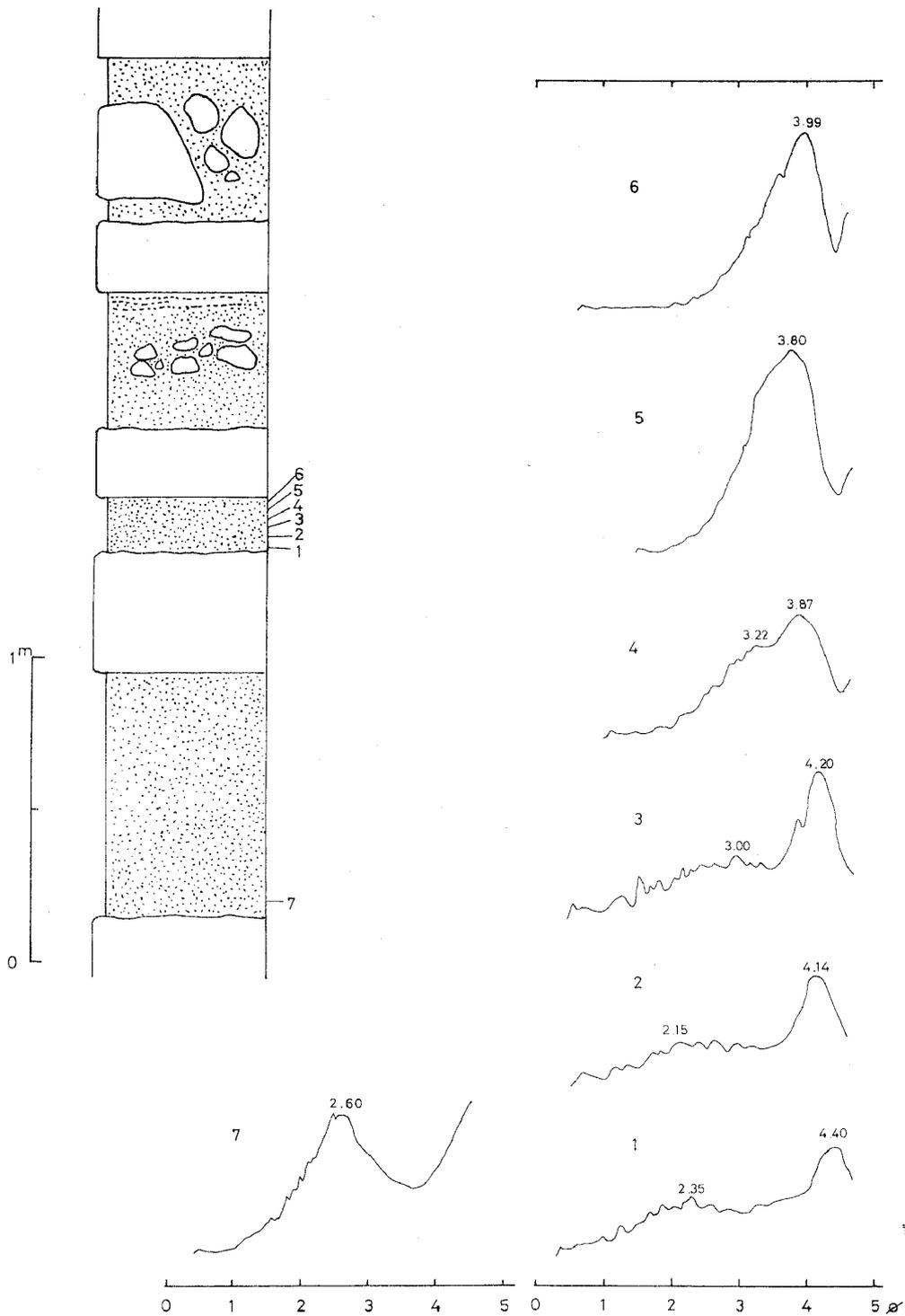


Fig. 4 Frequency curves of grain-size in sandstone alternating with siltstone.

ribution in the sandstone layers less than 100 cm thick, however, varies in parts, changing gradually from the lower part to the upper part (Fig. 4). That type of the pattern has not so far been found among the modern sand of Mutsu Bay and the northern part of the Pacific coast of Japan. The results of sedimentological analysis suggest that the sandstone was deposited under a special hydrodynamic condition. Meanwhile the surface of the

siltstone in the alternation is usually sculptured by the overlying sandstone and soft siltstone blocks are often included in the sandstone. It is obvious that the sands were once deposited in the neritic zone and then flowed down as so-called turbidity currents, during which courses the underlying siltstone was sculptured to be incorporated as soft siltstone blocks and as silt particles into the sand flow. Their deposition are instantaneous on the sea bottom, interrupting the deposition of the siltstones accumulated on the horizontal plane at a nearly constant rate of sedimentation under a calm condition.

The rate of sedimentation at the horizon near the boundary between the Matuyama and Brunhes epochs in the Kokumoto Formation was estimated in average to be 57 cm/1000 years (Nakagawa *et al.*, 1969). If the sandstones were formed instantaneously as is suggested above, the siltstone must have accumulated at a much slower rate. The ratio of the total thickness of the sandstones to siltstones in the sequence is 1.00:1.09 from the columnar section along the Yoro Valley (330–690 m by Mitsunashi *et al.*, 1961). The rate of sedimentation for siltstone in the sequence could be calculated at 29.7 cm/1000 years  $\left(=57 \times \frac{1.09}{1.09+1.00}\right)$  on the assumptions that (i) the rate of sedimentation for siltstone was constant, and (ii) the sandstone in the sequence was deposited instantaneously.

A rate of sedimentation for the siltstone at 29.7 cm/1000 years is larger than, for instance 0.75–1.13 cm/1000 years for red clay (Ninkovich *et al.*, 1966) and 1.98 cm/1000 years for *Globigerina* ooze (Oba, 1969) in the deep sea sediments. If this figure is correct, the 3.5 cm thick test piece measured in the present work represents an average period of 118 years, a duration short enough for examination of secular variation of the earth's magnetic field.

#### SAMPLES

It has already been known by the work by Nakagawa *et al.* (1969) on the basis of the measurements made for samples taken at 25 m stratigraphic intervals, that the boundary between the Matuyama and Brunhes Magnetic Polarity epochs lies in the middle part of the Kokumoto Formation: the interval over the boundary can be narrowed down to the levels between 170–200 m of the Kokumoto Formation in the columnar section along the Yoro Valley presented by Mitsunashi *et al.* (1959). This stratigraphic interval supposed to include the boundary was traced in the field by means of a set of key tuffs to the sequence showing similar lithology along a tributary of the Yoro Valley in the valley of Yanagawa (Fig. 5), where the condition for detailed continuous sampling is best suited.

Since the present study involves the correlation of results obtained by various lines of studies, a control is necessary as to the quantity and relative position of each sample used in laboratory studies and measurements. Furthermore, an oriented sample normal to the bedding plane is required for the paleomagnetic measurements. For this reason samples were taken in the field by directly coring the rocks at the outcrops. The Macand-rill with Mac-35A type engine of McCulloch Co. was used with two kinds of bits: a diamond bit of 35 mm in inner diameter and a metal-crown bit of 150 mm in inner diameter with core tubes 350 mm and 200 mm in length, respectively. In both cores water for sweeping the cuttings was supplied by means of water-swivel. The fresh part of the sedimentary rocks was cored vertically to their bedding plane.

The sampling from the Yanagawa section was made in three steps. Firstly, siltstone samples were taken at every one meter through the stratigraphic interval, 40 m thick (Fig. 5). The samples from KM 71–97 represent in an ascending order a 20 m thick sequence below the supposed boundary traced from the one found by the previous work (Nakagawa *et al.*, 1969); those from KM 11–23 represents, in an ascending order a 20 m

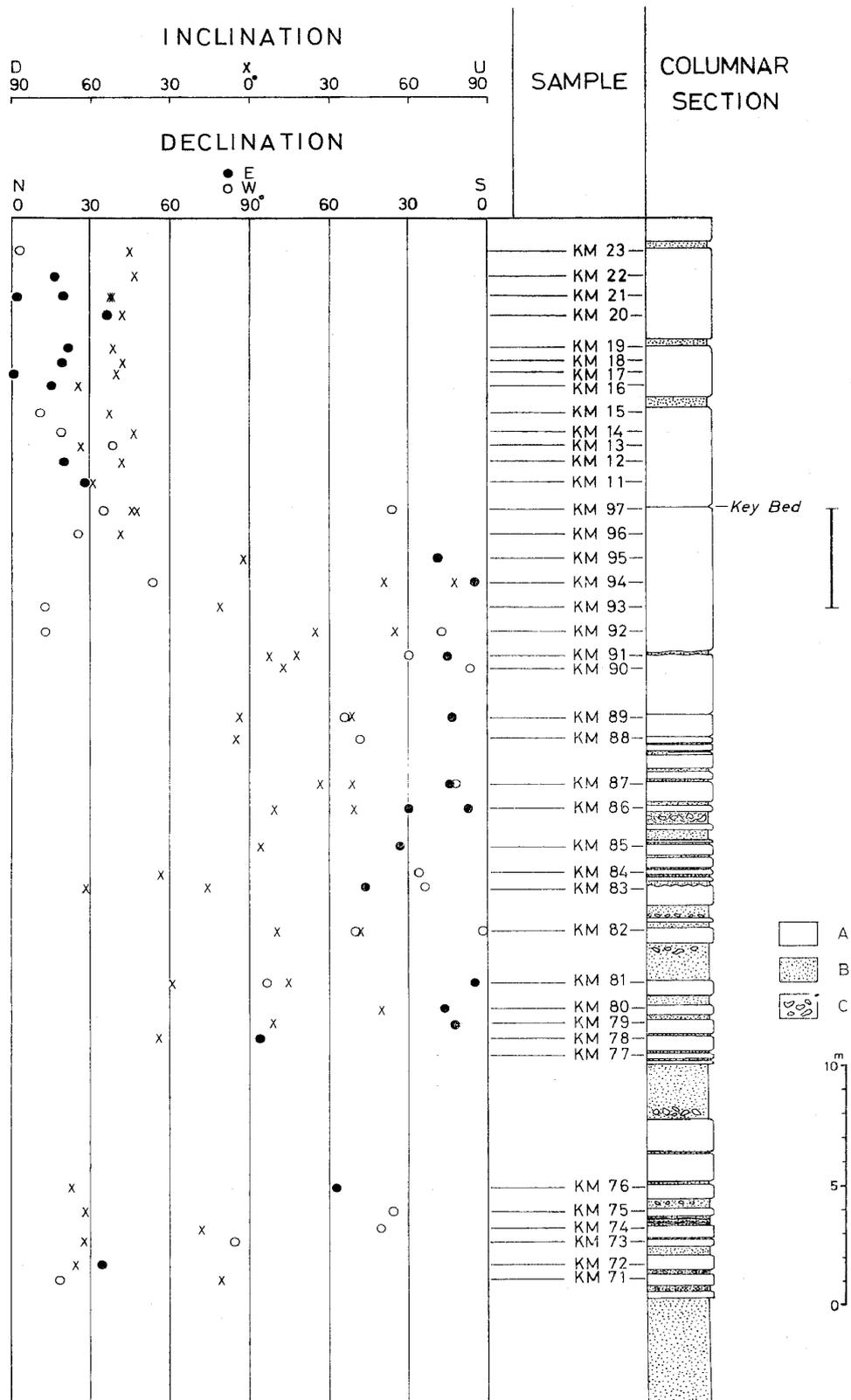


Fig. 5 Lithology and the results of paleomagnetic measurements from the Yanagawa section.  
 A: siltstone, B: sandstone, C: siltstone gravel

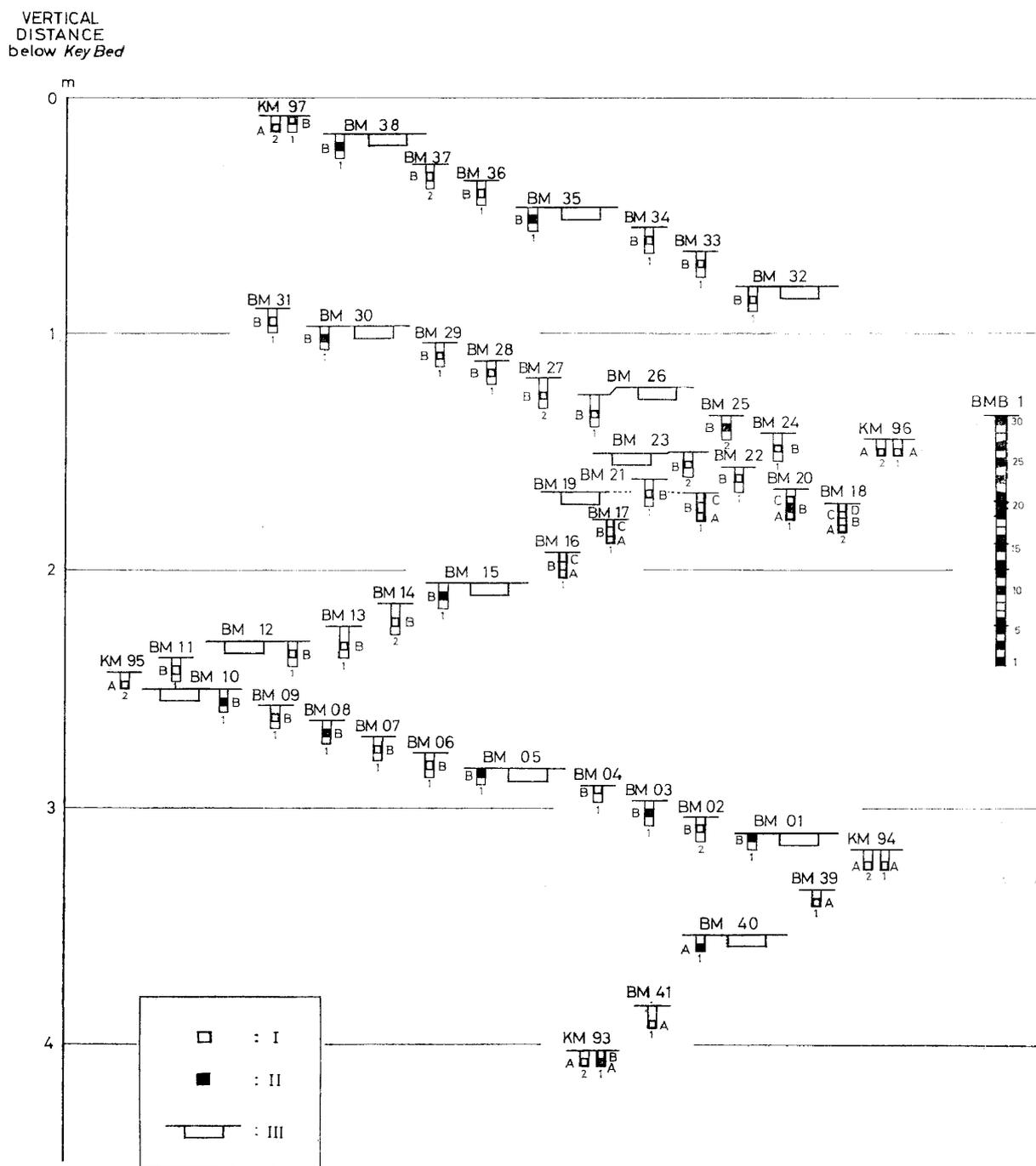


Fig. 6 The position of samples taken for the detailed study of the Matuyama-Brunhes boundary.  
 I: sample for paleomagnetic and sedimentological study  
 II: sample for paleomagnetic, sedimentological and foraminiferal study  
 III: sample for study of oxygen isotope

thick sequence above the supposed Matuyama and Brunhes boundary. At each stations, two cores were taken normal to the bedding plane, diameter and length being 35 mm and 120 mm, respectively. By the paleomagnetic measurements of these samples, the boundary was narrowed down further to exist between KM 95 and KM 96.

For the second steps for more detailed sampling about the boundary, samples were

taken at every 10 cm (BM 01-38) from the interval between KM 93 and KM 97 (Fig. 6). They consist of three cores, two cores of 35 mm in diameters and 120 mm in length, and another core of 150 mm in diameter and 50 mm length, taken at each site.

Moreover, the just boundary was crossed by a longer core (100 cm in length, 35 mm in diameter) taken to be continuous at the site (a series BMB 101-130).

Through these processes, the change of polarities representing the Matuyama and Brunhes boundary in the Yanagawa section was found, its position being 150 cm stratigraphically below the key tuff, a fine tuff layer 10 cm thick.

The samples were used in the following ways. Paleomagnetic measurements were made on 35 mm thick cylinders cut from 35 mm diameter cores. After the measurement, the test piece was cut longitudinally into three pieces; one half was for preservation, and one quarter for the foraminiferal studies. The samples taken by a 150 mm bit were used for collecting samples large enough to yield sufficient foraminiferal tests for oxygen isotope analysis.

#### EARTH'S MAGNETIC FIELD DURING REVERSAL

The method of measurement of DRM in the present work is almost the same as that described in Nakagawa *et al.* (1971). After the partial demagnetization, the direction and intensity of the remanent magnetization was measured at every 30° around three axes of the sample with the astatic-type magnetometer of the sensitivity of  $5 \times 10^{-8}$  Oe/mm. Each specimen was demagnetized in the alternation field of 90 Oersted to remove the unstable secondary components. The demagnetization was achieved with an apparatus similar to that described by Doell and Cox (1967), consisting of a tumbler rotating within an alternating field coil in nonmagnetic space provided by two pairs of Helmholtz coils. All numerical values obtained from the measurement were processed by an electronic computer, NEAC 2200, Model 500, at the Computer Center of the Tohoku University (see Appendix). Numerical analysis of the data indicates that the measured values can be well approximated by a magnetic dipole so that the direction and intensity of the remanent magnetization thus determined are sufficiently reliable.

The result of measurements for stratigraphic intervals of 1 m (Fig. 5) shows that northward declinations and downward inclinations of DRM prevail in the part above the key tuff indicating that they are of normal magnetic polarity. Below the key tuff, the declinations were southward and the inclinations upward, therefore they are of reversed magnetic polarity (Table 2).

The results of measurements for the sequences of 4m below the key tuff (Fig. 7, Table 3) show that the declination changed abruptly from southward to northward at a level of 170 cm below the key tuff; the inclination changed gradually from 40° upward to 40° downward between 240 cm below the key tuff and 100 cm below the key tuff. The inclination moved through the horizontal at 200 cm to 90° downward at 170 cm below the key tuff. The positions of virtual magnetic pole are calculated from these values and are plotted in Fig. 5. It seems to show that the virtual magnetic north pole migrated from the geographic south pole to the north pole along the meridian of 120° E.

The ratio of the intensity of DRM after AC-demagnetization by 90 Oe,  $J_n$ , to that of isothermal remanent magnetization after application of a direct field of 5000 Oe,  $J_R$ , is measure and plotted in Fig. 8. Although the intensity of the depositional remanent magnetization is not necessarily proportional to the intensity of the ambient field (Nagata, 1961), the values  $J_n/J_R$  may still be useful as an indication of the paleomagnetic intensity. Minima of  $J_n/J_R$  were recognized at the horizons 400, 235, 180 and 10 cm below the key tuff. The minimal ratios at the horizon 235 cm below the key tuff occurred at the beginning of

Table 2 Paleomagnetic measurements obtained for Yanagawa samples.

Sample	Declination (+ W, - E)	Inclination (+ U, - D)	Standard deviation	Intensity $10^3$ emu/cc
KM 23	N+ 3.0± 2.6	-45.3± 3.2	0.08, 0.12, 0.09	4.48
22	N-16.6± 6.2	-43.4± 6.2	0.10, 0.06, 0.09	3.46
21	N- 1.9± 0.6	-52.2± 5.0	0.04, 0.05, 0.23	6.05
	N-19.5± 5.3	-52.4± 5.0	0.05, 0.07, 0.07	7.15
20	N-36.1±10.8	-48.4± 8.2	0.11, 0.11, 0.10	7.85
19	N-21.4± 6.1	-51.7± 1.9	0.11, 0.09, 0.24	5.59
18	N-19.3± 8.6	-47.9± 8.3	0.07, 0.05, 0.06	4.57
17	N- 0.8± 7.5	-50.3±11.5	0.35, 0.07, 0.06	3.34
16	N-15.2±15.5	-64.8±10.6	0.06, 0.15, 0.11	3.65
15	N+10.7± 1.1	-52.6± 6.3	0.07, 0.11, 0.15	3.59
14	N+18.9± 3.5	-43.4± 1.7	0.06, 0.14, 0.09	3.50
13	N+38.8± 1.0	-63.1± 3.3	0.08, 0.06, 0.09	2.34
12	N-20.0± 7.2	-48.1± 7.1	0.16, 0.05, 0.16	3.46
11	N-28.1± 8.9	-59.5± 4.4	0.08, 0.09, 0.07	4.90
97	S+36.1±24.7	-44.7±37.7	0.18, 0.55, 0.23	0.73
	N+34.7± 0.2	-42.4±19.3	0.34, 0.16, 0.63	0.92
96	N+25.4±15.7	-48.7± 1.2	0.14, 0.16, 0.14	1.41
95	S-18.9± 8.0	- 2.4± 1.4	0.10, 0.18, 0.03	2.82
94	S- 4.8± 7.2	+50.9± 5.7	0.07, 0.12, 0.15	1.80
	N+53.3±11.4	+76.9± 3.3	0.17, 0.13, 0.78	2.33
93	N+12.6± 5.4	-11.5±29.7	0.21, 0.32, 0.17	0.96
92	N+12.7±18.6	+35.5± 4.1	0.16, 0.35, 0.11	2.60
	S+17.2±18.0	+24.7± 5.2	0.16, 0.17, 0.04	4.16
91	S-15.7±25.0	+ 7.2±27.9	0.16, 0.32, 0.09	1.26
	S+29.5± 6.7	+17.7± 3.0	0.07, 0.06, 0.06	4.48
90	S+ 6.4± 0.5	+12.3± 9.8	0.22, 0.07, 0.11	2.86
89	S+52.8±14.9	- 3.9±10.7	0.32, 0.14, 0.16	2.28
	S-13.5± 8.1	+38.1±24.1	0.13, 0.09, 0.04	2.37
88	S+48.6±33.2	- 4.9± 1.9	0.39, 0.53, 0.68	0.54
87	S+12.5±16.7	+38.9±11.6	0.08, 0.22, 0.15	1.44
	S-14.8±39.1	+27.2±13.3	0.14, 0.10, 0.03	2.82
86	S-29.8± 0.2	+ 9.1± 6.8	0.18, 0.15, 0.08	1.82
	S- 7.7± 1.0	+39.6±14.2	0.10, 0.13, 0.17	3.34
85	S-33.6±68.2	+ 3.8±11.9	0.29, 0.82, 0.06	1.85
84	S+26.2± 4.6	-33.5± 2.7	1.34, 0.10, 0.17	1.20
83	S-45.8±25.0	-61.9± 8.8	0.20, 0.33, 0.14	1.48
	S+23.8±19.3	-15.8± 4.6	1.25, 0.25, 0.07	2.07
82	S-49.5±10.5	+10.4±22.5	0.33, 0.34, 0.12	0.83
	S+ 1.8±93.2	+40.2± 0.5	0.45, 0.03, 0.11	1.18
81	S- 4.7±87.3	-29.2±39.7	1.53, 0.43, 0.27	0.20
	S+83.4±98.9	+14.2±46.6	0.24, 0.38, 0.11	1.37
80	S-16.3± 0.8	+49.8±12.4	0.06, 0.16, 0.20	2.30
79	S-12.5± 6.8	+ 8.8± 9.9	0.15, 0.31, 0.07	2.94
78	S-86.7± 1.7	-34.6±45.1	0.20, 0.25, 0.18	3.68
76	S-57.4±39.2	-67.1±22.7	0.10, 0.13, 0.12	3.51
75	S-35.6±83.2	-62.1±14.6	0.21, 0.38, 0.12	3.54
74	S+40.6±22.3	-18.1± 5.1	0.76, 0.37, 0.08	2.15
73	N+84.7±19.4	-62.6± 5.1	0.25, 0.26, 0.14	2.16
72	N-34.5±36.3	-65.3±20.6	0.09, 0.18, 1.66	2.26
71	N+18.3± 3.9	-10.8± 5.6	0.59, 0.41, 0.09	1.87

the magnetic field reversal shown by inclinations. During the geomagnetic field reversal was recognized between the horizons 240 cm and 100 cm below the key tuff, the intensity apparently remained in smaller values. As Fig. 7 indicates in a general tendency, the intensity of the paleomagnetic field oscillated periodically, with a period represented by 200 cm in thickness of sediments. The maximum values of the  $J_n/J_R$  are about three times larger than the minimum values. During a polarity change, the intensity of the field decreased, the magnetic field, however, never vanished as was suggested by the oriented DRM (Fig. 8).

At the two horizons of minimal intensity, 400 and 100 cm below the key tuff, the

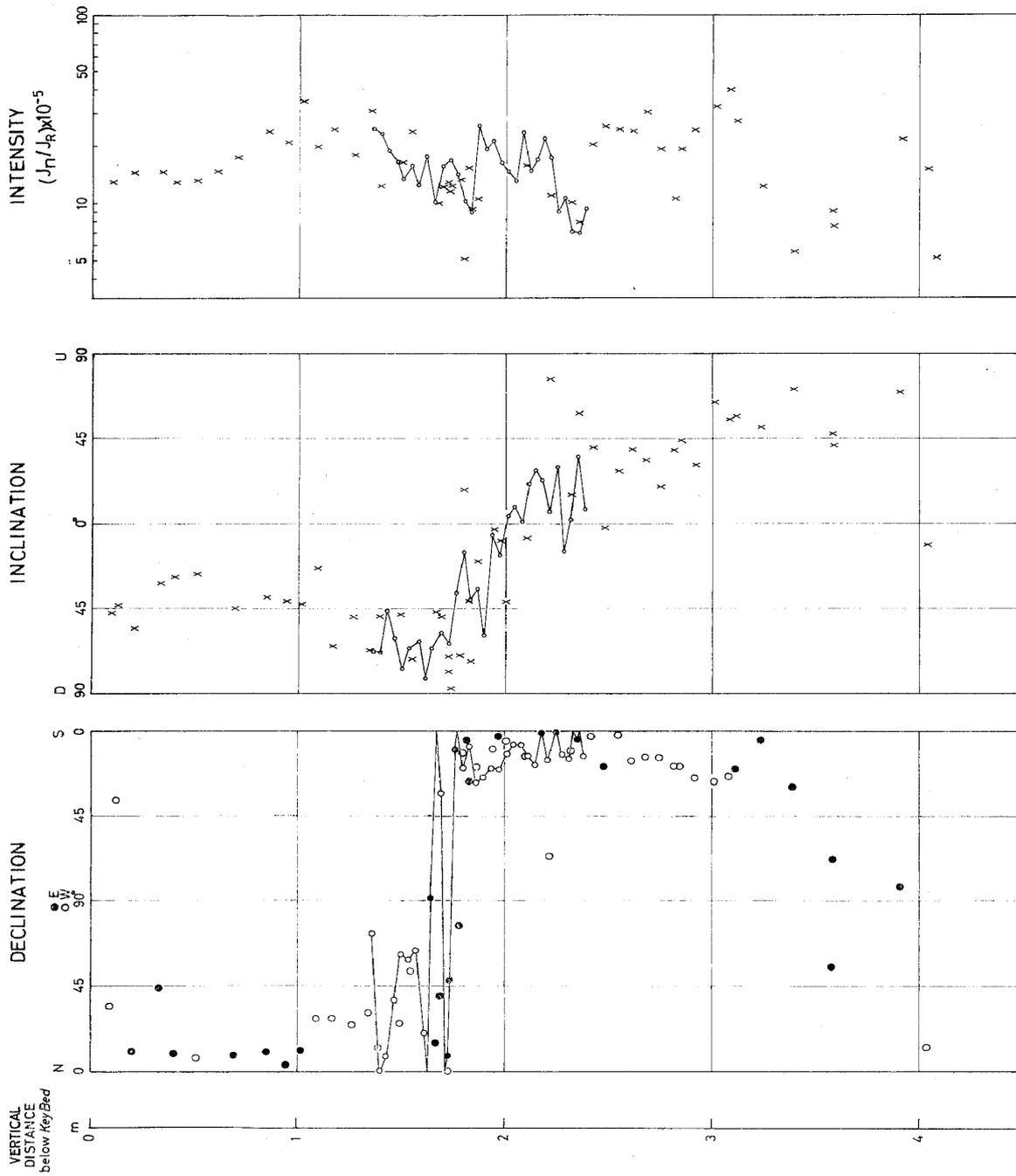


Fig. 7 Results of paleomagnetic measurements.

Each plot represents a sample; those taken from a single continuous core are connected by lines.  $J_n/J_R$  is a ratio of the intensity of DRM after AC-demagnetization by 90 Oe to the intensity of isothermal remanent magnetization after application of a direct field of 5000 Oe.

Table 3a Paleomagnetic measurements obtained for the Matuyama-Brunhes boundary.

Sample	Declination (+ W, - E)	Inclination (+ U, - D)	Standard deviation	Intensity $\times 10^{-3}$ emu/cc	$J_n/J_R$ $\times 10^{-5}$
HM971B	N+34.7± 0.2	-42.4±19.3	0.34, 0.16, 0.63	0.92	12.8
972A	S+36.1±24.7	-44.7±37.7	0.18, 0.55, 0.23	0.73	
BM361B	N-10.6±17.3	-55.2± 8.6	0.12, 0.12, 0.18	1.07	14.4
372B	N-44.0±33.0	-31.9± 6.9	0.29, 0.22, 0.09	1.22	14.6
361B	N- 9.4±17.6	-28.7± 8.3	0.35, 0.28, 0.11	1.13	12.8
351B	N+ 7.1±10.2	-26.8± 4.1	0.27, 0.09, 0.19	1.12	13.1
331B	N- 8.4±17.2	-45.6±12.1	0.22, 0.13, 0.14	1.66	17.3
321B	N-10.3± 5.4	-39.2± 7.2	0.09, 0.14, 0.12	3.06	23.7
311B	N- 3.6±13.7	-41.4± 3.4	0.11, 0.14, 0.07	2.31	20.5
301B	N-10.9± 6.7	-42.8± 6.4	0.07, 0.07, 0.02	7.83	34.3
291B	N+28.0± 1.5	-23.9± 2.0	0.29, 0.21, 0.06	1.96	19.5
281B	N+27.9± 9.1	-65.1± 0.9	0.08, 0.10, 0.08	2.00	24.2
272B	N+24.8± 1.7	-49.7± 4.6	0.19, 0.15, 0.14	1.58	17.6
261B	N+30.9± 5.5	-67.4± 1.0	0.15, 0.16, 0.14	2.75	31.0
252B	N+12.6± 6.6	-49.4± 5.7	0.28, 0.10, 0.14	0.87	12.2
HM961A	N+25.4±15.7	-48.7± 1.2	0.14, 0.16, 0.14	1.41	16.4
BM232B	N+52.9±14.0	-72.2± 5.3	0.13, 0.16, 0.13	2.71	23.9
211B	N-15.1±18.6	-47.0± 5.4	0.22, 0.25, 0.06	1.51	9.7
191C	N-40.1±34.9	-49.9±14.3	0.24, 0.23, 0.27	1.61	12.0
201B	N+ 0.2±18.9	-70.8± 7.7	0.34, 0.20, 0.11	1.85	12.8
132D	N+38.8±24.9	-79.1±19.2	0.10, 0.31, 0.39	1.69	11.5
191B	N-48.7±66.5	-87.4± 3.8	0.23, 0.24, 0.51	1.32	12.4
191A	N-77.0±58.7	-69.7±15.6	0.17, 0.15, 0.42	1.24	13.3
191C	S+11.5± 2.1	+17.8±14.5	0.22, 0.27, 0.12	0.50	5.03
132A	S- 4.4± 0.8	-41.6± 1.0	0.14, 0.13, 0.10	1.42	15.3
171B	S-26.2±47.9	-73.6± 8.8	0.21, 0.09, 0.11	0.91	9.3
171A	S+18.5±12.6	-20.3± 8.1	0.18, 0.19, 0.13	0.91	10.4
161A	S+ 5.8± 0.3	-41.7± 8.9	0.16, 0.22, 0.09	0.83	
151B	S+12.8± 9.0	- 8.0± 4.2	0.10, 0.23, 0.03	1.89	15.8
142B	S+65.9±95.3	+75.6± 8.7	0.23, 0.18, 0.09	1.39	10.9
131B	S+10.6±16.8	+15.1± 5.1	1.54, 0.17, 0.03	0.93	10.1
121B	S- 2.3±14.4	+57.8± 2.0	0.49, 0.11, 0.35	0.89	7.9
111B	S+ 2.4± 1.5	+40.0± 6.2	0.34, 0.22, 0.03	1.70	20.0
HM952A	S-18.9± 8.0	- 2.4± 1.4	0.11, 0.18, 0.03	2.82	25.3
HM101B	S+ 2.0± 4.6	+27.7± 2.2	0.10, 0.11, 0.09	2.83	24.5
091B	S+15.4± 2.6	+39.0± 2.8	0.13, 0.07, 0.08	2.77	23.9
081B	S+13.9±16.2	+33.4± 9.4	0.15, 0.04, 0.24	3.72	30.4
071B	S+13.9± 1.7	+19.5± 1.6	0.10, 0.05, 0.03	3.18	19.1
061B	S+18.6±22.2	+38.8±13.1	0.15, 0.14, 0.07	2.04	10.6
051B	S+13.4± 7.1	+44.4± 7.2	0.11, 0.06, 0.10	3.27	19.2
041B	S+25.1± 5.6	+30.9± 2.6	0.25, 0.09, 0.06	3.45	24.4
031B	S+26.7±19.7	+64.1±10.4	0.13, 0.10, 0.32	4.85	32.8
022B	S+24.0± 8.4	+54.8± 3.7	0.12, 0.17, 0.22	3.05	40.2
011B	S-19.9± 7.2	+56.5± 1.5	0.13, 0.10, 0.15	3.67	27.0
HM941A	S- 4.8± 7.2	+51.0± 5.7	0.07, 0.12, 0.15	1.80	12.4
BM991A	S-29.5± 6.3	+71.4± 5.2	0.26, 0.28, 0.14	0.85	5.57
401A	S-68.2±98.3	+47.5± 0.7	0.38, 0.12, 0.14	1.76	9.2
402A	N-54.9±18.1	+41.5± 5.5	0.24, 0.16, 0.14	1.46	7.68
411A	S-82.9± 0.2	+69.8± 0.3	0.17, 0.09, 0.63	3.42	22.2
HM931B	N+12.6± 5.4	-11.5±29.7	0.21, 0.32, 0.17	0.96	5.19

declination varied greatly, though the inclination remained approximately the same. Such behavior of geomagnetic field has once been reported by Petorova (1965) as the behavior of "step I". The position of the virtual magnetic pole calculated from the horizons of the minimum intensity drifted along the meridians of 180°, 120° and 90° of east longitude at horizons 400 cm, 200 cm, and 10 cm below the key tuff, respectively. The drift of the virtual pole may be caused by the increase of the proportion of non-dipole moment in the total geomagnetic field caused by the main dipole moment decreasing at those horizons.

The duration of the magnetic field reversal can be calculated under the assumption that the rate of sedimentation is uniform at a magnetically dated stratigraphic interval.

Table 3b Paleomagnetic measurements obtained for the Matuyama-Brunhes boundary. (continuous core)

Sample	Declination (+ W, - E)	Inclination (+ U, - D)	Standard deviation	Intensity $\times 10^{-3}$ emu/cc	$J_n/J_R$ $\times 10^{-3}$
EMBL30	N+72.3± 7.4	-68.7± 7.7	0.13, 0.27, 0.37	2.08	24.6
129	N+ 0.3± 2.7	-68.8± 5.1	0.17, 0.07, 0.16	1.81	23.2
123	N+ 7.9± 0.2	-46.5±12.4	0.25, 0.23, 0.19	1.72	18.7
127	N+37.5± 5.8	-60.9± 1.1	0.28, 0.34, 0.23	1.47	16.4
126	N+61.8± 2.2	-77.3± 1.9	0.18, 0.16, 0.35	1.27	13.3
125	N+58.9± 6.5	-66.4± 5.2	0.14, 0.11, 0.16	1.72	15.5
124	N+64.1± 4.9	-63.1± 0.5	0.20, 0.12, 0.15	1.29	12.3
123	N+20.0± 7.3	-82.2± 7.3	0.13, 0.19, 0.19	1.73	17.5
122	S-85.7± 5.6	-66.6± 9.5	0.31, 0.19, 0.20	1.23	10.0
121	S+32.8±29.3	-57.9±12.2	0.16, 0.14, 0.20	2.65	15.7
120	K- 7.9± 9.9	-63.9± 4.3	0.10, 0.10, 0.13	2.90	16.5
119	S-10.0±19.6	-34.7± 2.2	0.30, 0.15, 0.19	1.52	14.0
118	S+19.0±17.2	-15.6± 2.0	0.43, 0.30, 0.13	0.85	10.1
117	S+ 8.5±22.2	-41.0± 2.4	0.56, 0.19, 0.35	0.81	8.8
116	S+27.5±17.8	-35.4± 4.0	0.18, 0.21, 0.16	3.03	25.6
115	S+24.9±13.4	-60.4± 4.0	0.15, 0.22, 0.24	2.53	19.1
114	S+19.8± 6.4	- 6.6± 2.4	0.30, 0.31, 0.11	2.06	21.0
113	S+20.6±17.9	-17.4± 3.9	0.25, 0.17, 0.11	1.64	16.2
112	S+12.1±12.1	+ 3.3± 5.0	0.56, 0.12, 0.08	1.46	14.6
111	S+ 7.2± 5.8	+ 8.3± 3.7	0.40, 0.12, 0.10	1.65	13.0
110	S+ 7.4±11.2	+ 0.1± 3.4	0.40, 0.13, 0.04	3.01	23.8
109	S+13.1± 1.2	+20.3± 2.8	0.24, 0.09, 0.05	2.70	14.6
108	S+18.2± 2.5	+27.5± 2.2	0.26, 0.03, 0.05	2.26	16.9
107	S- 0.2± 2.1	+22.4± 4.6	0.28, 0.05, 0.11	2.07	21.9
106	S+15.4±20.0	+ 5.9± 5.5	0.33, 0.17, 0.06	1.69	17.2
105	S- 0.5± 2.6	+29.2± 7.8	0.23, 0.14, 0.12	0.98	8.9
104	S+12.5±27.4	-15.3± 1.3	0.56, 0.24, 0.35	1.06	10.5
103	S+14.2±29.8	+ 1.6± 8.0	0.24, 0.29, 0.21	0.73	6.9
102	S- 4.3± 3.5	+35.0± 8.1	0.33, 0.20, 0.19	0.75	6.9
101	S+13.2±19.5	+ 7.2±28.4	0.29, 0.18, 0.29	1.07	9.2

The time required for the polarity transition was 4700 years, because it corresponds to the sedimentary record of 140 cm thick sediments (an interval between 240–100 cm below the key tuff). This value is in fairly good agreement with the period of 4600 years independently estimated by Cox *et al.* (1967) on the basis of statistical treatments of the polarity record in igneous rocks dated by the K-Ar method. The period of the geomagnetic intensity fluctuation, corresponding to the 200 cm thick sediments, is estimated to be 6700 years. This figure is also almost consistent with the 8000 years period of intensity fluctuations estimated from the archeomagnetic data (Smith, 1967a,b, 1968).

Exact correlation of the world wide time plane will be possible with an error within about several hundreds years by tracing the positions of the virtual magnetic pole which drifted along the meridian during the reversal (Fig. 8). It will be also possible to trace the world wide time planes with an error less than 1000 years by the measurements of continuous samples taken from a stratigraphically controlled section to utilize the intensity fluctuations of the paleomagnetic field around the horizon of the reversal. If that the period of the intensity fluctuations of the paleomagnetic field was 7000–8000 years throughout geologic time is proved reliable, this may warrant the estimation of the rate of sedimentation from the measurement of the intensity fluctuations and the sedimentary thickness.

#### SEDIMENTOLOGICAL ANALYSIS

As previously mentioned, the sedimentological analysis was made on a quarter cut of the cored sample used for the paleomagnetic measurement. Altogether 72 samples

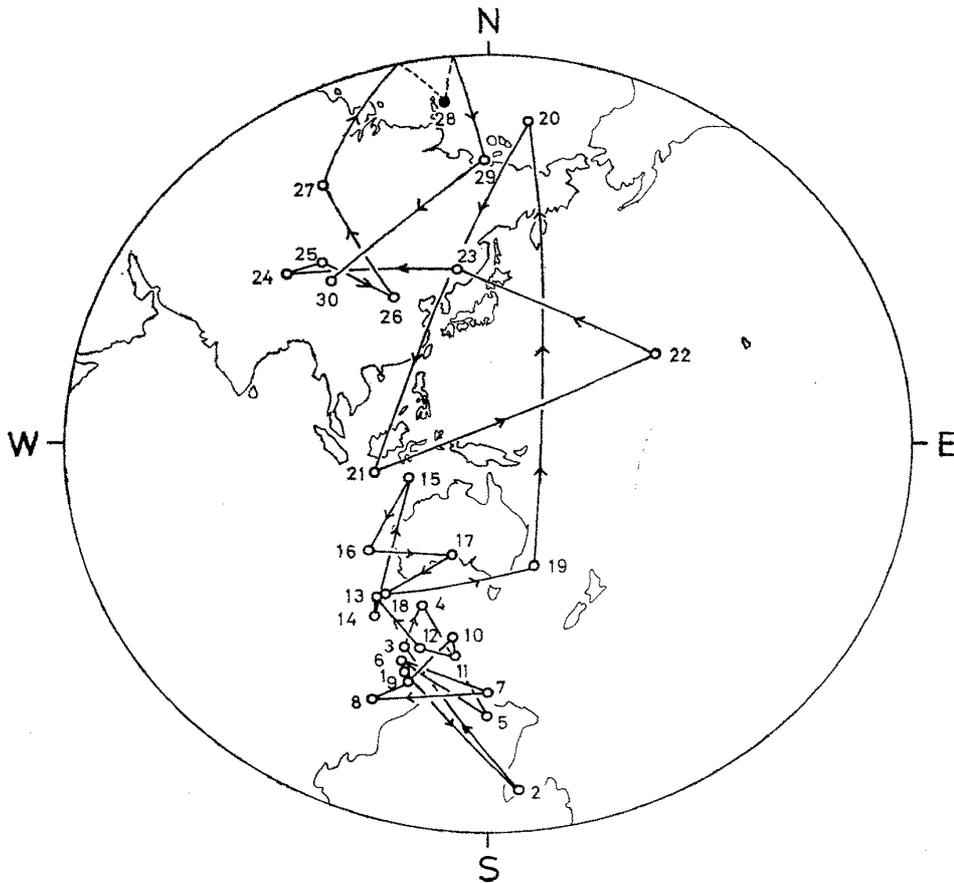


Fig. 8 Drift of virtual magnetic north pole during the reversal at the Matuyama-Brunhes boundary. The positions of the virtual pole are calculated from the data which are represented by points connected by line in Fig. 7.

have been analysed; their stratigraphic positions are shown in Fig. 5.

After soaked in water for one hour, the rock sample was disintegrated in water by a speed controlled mixer. The mud content (fraction of the grains finer than  $4.0\phi$ ) and frequency curve for the fraction coarser than  $5.0\phi$  were obtained by the automatic grain-size analyser (Niitsuma, 1971).

The mud contents of the samples are rather uniform ranging from 70 to 85 percent, with two weak maxima at 65 cm and 260 cm below the key tuff (Fig. 9).

The grain-size distributions were found to be more or less similar in all samples, in that the fraction coarser than  $3.0\phi$  is small, and the fine fraction predominates. Sediments having a similar grain size distribution have been found in the studies of modern sediments along the Japanese coasts (Niitsuma and Mekata, 1971 a,b). They are found in areas where bottom conditions are judged to be calm hydrographically, although the depths of the water varies greatly. By analogy, the sediments from the horizons concerned in this study are interpreted to have been formed in a similarly calm condition. The comparison of the frequency curves obtained for each sample shows that they can be grouped into four types, representatives of which are shown in Fig. 10. The curves shown that the differences in composition exist in the fraction around  $4\phi$  in size; it appears that they are caused mainly by the removal of the grains around  $4\phi$  in size. It is considered that the sediments of type IV were the most affected by the removal while those of type I were unaffected.

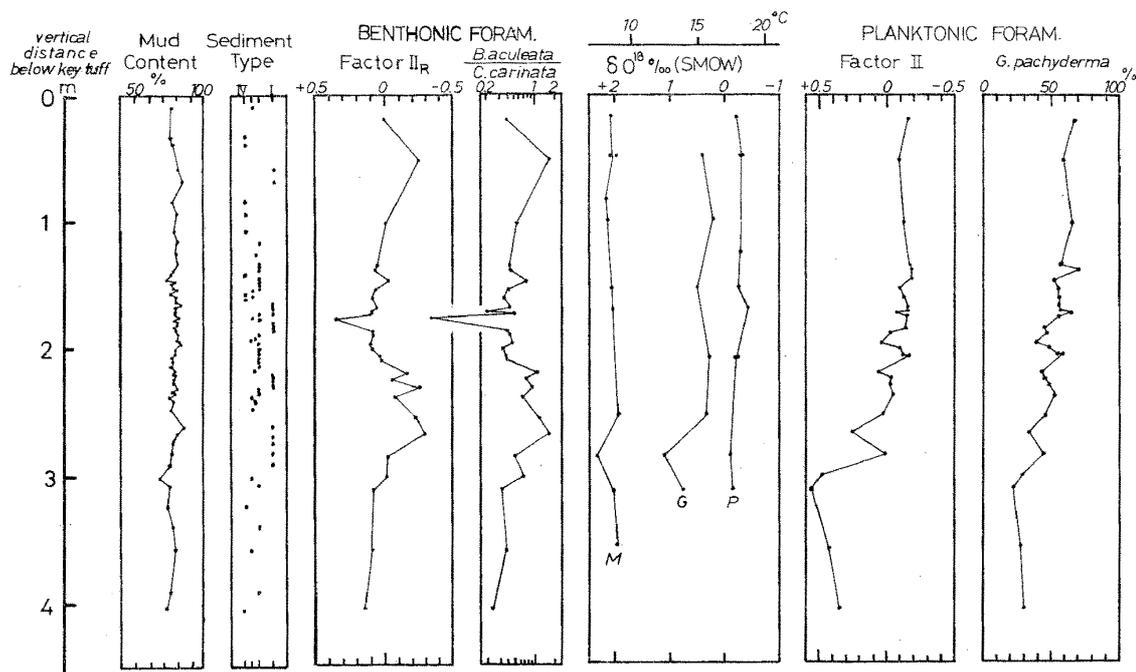


Fig. 9 Results of the grain-size analysis, paleotemperature measurement and the faunal analyses of Foraminifera. The samples dealt with are the same as those subjected to the paleomagnetic measurements (Fig. 4).

M: *Melonis barleeanus* G: *Globigerina pachyderma* P: *Pulleniatina obliquiloculata*

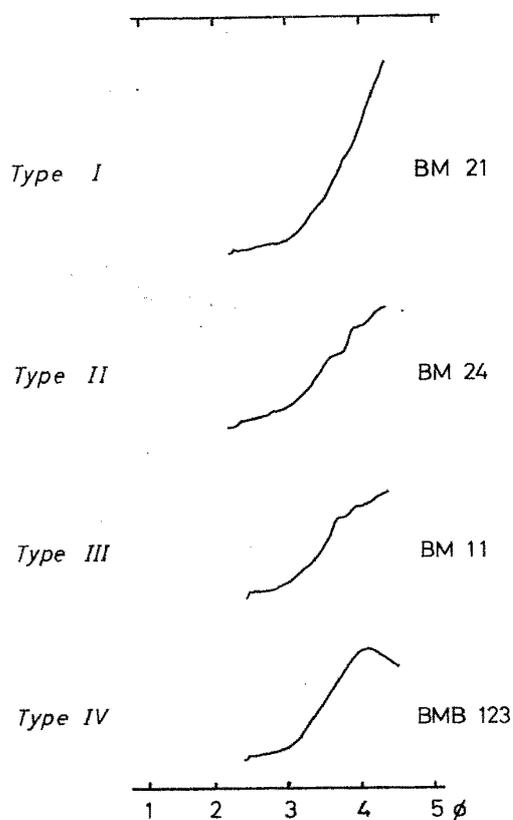


Fig. 10 Frequency curves of grain-sizes obtained by the automatic grain-size analyser. They represent four types.

In the present section the samples that fall in the type I are found from the horizons representing the five stratigraphic intervals as follow; 295 to 255, 213 to 210, 188 to 179, 173 to 163, and 80 to 50, all in cm below the key tuff (Fig. 9). These intervals which represent the calmest conditions are separated by other types of sediment. Although mud contents are rather uniform throughout the section, a tendency exists that they are higher in the above mentioned five intervals represented by type I. Of the five intervals, two, the uppermost and the lowermost correspond to the two maxima in mud contents. On the basis of these considerations it may be said that the water was deeper during the accumulation of the said two intervals than during the deposition of the rest of the section.

#### PALEOTEMPERATURE ANALYSIS

The purpose of the present study is to determine the changes in temperature of sea-water as related to the reversal of the earth's magnetic field. The temperature of the sea water was determined by measuring the oxygen 18 content of the tests of Foraminifera which were abundant in the siltstones dealt with in the present study.

For the biogeneous calcium carbonate to be suitable for the purpose, the following requirements must be met. They are:

- 1) the biogeneous calcium carbonate must have been formed in the oxygen isotopic equilibrium with sea water,
- 2) the material must be contained in rocks in abundance so that one can gather a sufficient amount (ca. 20 mg) from the samples from every horizons taken at short stratigraphic intervals throughout the section studied,
- 3) the materials are tests of a single species throughout the section by which the influence due to the difference in the ecological behavior due to species can be eliminated,
- 4) the material is composed of sufficient numbers of monospecies specimens so that the individual difference in isotopic ratio can be averaged,
- 5) the ecological behavior of the species whose tests are used should be known, and
- 6) the oxygen isotope ratio of the tests must have remained unchanges since the biogeneous calcium carbonate was formed.

Among the many species of Foraminifera found in the section, the planktonic Foraminifera *Pulleniatina obliquiloculata*, *Globigerina pachyderma*, and the benthonic Foraminifera *Melonis barleeanus* were chosen for the oxygen isotope analysis.

The samples were collected at every 20 to 50 cm intervals vertically against the bedding plane between the key tuff and 4 m below it using a core sampler of 150 mm in diameter. The thickness of one sample was 5 cm and its weight was approximately 1000 g. After soaked in water for 24 hours, the rock sample was disintegrated in a speed controlled mixer. The sample was wet sieved through 42 and 80 meshes. The fraction between 42 and 80 meshes in size was dried and subjected to the separation into magnetic and nonmagnetic fraction by means of the Frantz isodynamic separator with the side tilt 30°, backward tilt 10°, and magnet charged with 1.0 A current. The magnetic fraction contains rock fragments and heavy mineral grains, whereas in the nonmagnetic fraction the remains of Foraminifera, Radiolaria, and the skeltons of other fossils are concentrated.

The tests of Foraminifera were picked up from the nonmagnetic fractions under a binocular microscope. Most of the chambers of the foraminiferal test were filled with authogenic pyrite. The foraminiferal tests of monospecies were pulverized with an agate mortar and an ultrasonic generator. After drying, the fragments of the tests were separated from pyrite with the Frantz isodynamic separator (side tilt 30°, backward tilt 0.5°, and charged current of magnet 1.75 A).

The pulverized monospecies material was purified by roasting to destroy the organic substances in the sample powder. The apparatus used was the same as that described by Epstein *et al.* (1953), Horibe and Oba (1969), and Oba (1969) that roasts the material on Pt boats at 470°C with a stream of helium flowing at about 0.4 cc/sec. for 30 minutes. After this treatment the sample was reacted with pure H<sub>3</sub>PO<sub>4</sub> to obtain CO<sub>2</sub> gass in a thermostat under the temperature precisely controlled at 25.00°C. The deviation  $\delta^{18}\text{O}$  ‰ of isotope ratio  $^{18}\text{O}/^{16}\text{O}$  in the CO<sub>2</sub> gass generated from each sample from that in the standard mean ocean water (SMOW) was measured by the Urey-Nier type mass spectrometer "PANDORA" (Horibe, 1966). The error in the oxygen isotope measurement mostly depends on the amount of the calcium carbonate sample; the error will be 0.028 ‰, when it weighs more than 12 mg, 0.079 ‰ when calcium carbonate is 5 mg, and 0.137 ‰ when it is 4 mg. Twelve milligrams of the sample represents approximately 1000 tests of *P. obliquiloculata*, 3000 tests of *G. pachyderma*, or 2000 tests of *M. barleeanus*.

Assuming that the oxygen isotopic composition of sea water in which the foraminiferal tests were formed was the same as that of SMOW, the paleotemperature can be determined from  $\delta^{18}\text{O}$  values of oxygen in the calcium carbonate tests by means of the scheme presented by Craig (1965).

The observed  $\delta^{18}\text{O}$  values (SMOW) ranged from -0.1 to -0.3 ‰ (from 17.3 to 18.1°C in paleotemperature) for the tests of *P. obliquiloculata*, from +0.2 to 1.1 ‰ (from 12.4 to 16.1°C) for *G. pachyderma*, and from +1.9 to +2.4 ‰ (from 7.2 to 9.2°C) for *M. barleeanus* (Table 4). The fact that these three species from the same sample gave quite different temperature from each other throughout the sequence appear to rule out the

Table 4 Paleotemperature measurements by Oxygen-isotope analysis.

Sample	species	$\delta^{18}\text{O}$ ‰ (SMOW)	Sens of measurement ‰/div.
BM 38	P	- 0.21	0.785
	M	+ 2.07	1.368
BM 35	P	- 0.35	0.785
	P	- 0.32	0.785
	G	+ 0.41	0.993
	M	+ 2.10	0.785
	M	+ 1.97	0.785
BM 32	M	+ 2.17	1.368
BM 30	G	+ 0.20	0.993
	M	+ 2.15	0.785
BM 26	P	- 0.29	0.785
BM 23	P	- 0.25	0.785
	G	+ 0.49	0.993
	M	+ 2.06	1.368
BM 19	P	- 0.44	1.368
	M	+ 2.03	0.785
BM 15	P	- 0.20	0.276
	P	- 0.23	0.276
	G	+ 0.27	0.993
BM 10	G	+ 0.38	0.993
	M	+ 1.93	0.785
BM 05	P	- 0.12	0.785
	G	+ 1.11	0.993
	M	+ 2.35	0.785
BM 01	P	- 0.13	0.785
	G	+ 0.77	0.993
BM 40	M	+ 2.06	0.785
	M	+ 1.98	0.785

P: Pulleniatina obliquiloculata  
G: Globigerina pachyderma  
M: Melonis barleeanus



possibility of the exchange of the oxygen isotope ratio of the tests by diagenetic effects.

The  $\delta^{18}$  value and paleotemperature at each stratigraphic horizon obtained for three foraminiferal species are shown in Fig. 9. No appreciable change is apparent from *P. obliquiloculata* and *M. barleeanus*. From *G. pachyderma*, however, the appreciable fluctuation of paleotemperature can be detected; it shows that the paleotemperature of sea-water in which *G. pachyderma* lived was 13.8°C at the horizon 313 cm below the key tuff, the lowest temperature 12.4°C at 286 cm below the key tuff and the paleotemperature rose to 15.4°C at the horizon 223 cm below the key tuff. Although the difference is small, it is of interest to point out that *M. barleeanus* showed almost a constant temperature of 9.2°C throughout the present section, has its minimum temperature (7.1°C) at the horizon 286 cm below the key tuff. At this horizon, 286 cm below the key tuff, *G. pachyderma* also showed lowest temperature (Fig. 9).

The constant difference of paleotemperature, ca. 10°C is found between the planktonic species, *P. obliquiloculata* and the benthonic species, *M. barleeanus*. It is considered that the difference of temperature was between the surface and bottom water column. Judging from the temperature gradient observed in the surface water column of the Pacific (Reid, 1965), the depth of water in which the sediments accumulated was estimated to be in the range of 200–300 meters.

#### CHANGE OF FORAMINIFERAL FAUNA

As mentioned previously the foraminiferal investigation was made on a quarter cut of the cored test piece used for the paleomagnetic measurements. The 29 samples were collected from the 4 m thick interval below the key tuff at stratigraphic intervals ranging from 3.5 to 50 cm (Fig. 5). After soaked in water for one hour, the rock sample was disintegrated in a water bath. After 15 minutes run in a controlled mixer at its slowest speed, the sample was wet sieved through 200 mesh and dried. The tests of Foraminifera were concentrated by making use of the Frantz isodynamic separator (side tilt 30°C, backward tilt 10°C, and magnet charged with 1.0 A current) and were picked up under a binocular microscope. In the present article, the faunal change of Foraminifera is analysed by the principal factor analysis (Harman, 1967; Imbrie *et al.*, 1962) and in terms of Motomura's Plane (Niitsuma, 1968). All calculations have been done with the electronic computer (NEAC 2200, Model 500) in the Computer Center of the Tohoku University (Niitsuma *et al.*, 1971).

##### 1) Faunal change of benthonic Foraminifera

In the present study, only the tests of Foraminifera coarser than 74 microns have been identified. At least 200 benthonic specimens have been identified for each sample, resulting in recording altogether 95 species in 50 genera from the samples. In the samples studied the species that occur in abundance are: *Epistominella naraensis* (Kuвано), *Cassidulina carinata* Silvestri, *Melonis barleeanus* (Williamson), *Bulimina aculeata* d'Orbigny, *Pseudoeponides japonicus* Uchio, *Cassidulina depressa* Asano and Nakamura, *Elphidium kusiroense* Asano. Their abundance point to the faunal characteristics of the zone 6 defined by Aoki (1963) in the Yoro Valley section (Table 5).

On the basis of relative frequencies of 70 species, factor analysis has been made to determine the relationship between benthonic foraminiferal faunas of 29 samples (Table 6). All loadings of the first factor are found positive and the loadings express the common component in faunas of all samples. The loadings of the other factors express the different component. The total loading of the first factor in all faunas is 91.0%, that of the second factor is 2.2% and that of the third factor is 1.85%. However, the first factor express the

Table 6 Data of Factor analysis of benthonic foraminiferal faunas.

Sample	FACTOR 1	FACTOR 2	FACTOR 3	FACTOR 4	FACTOR 5	FACTOR 6
KM 93	.95363	.10031	-.09732	-.15790	-.02521	-.02630
BM 40	.88757	.14182	.02039	-.11276	-.34265	-.02517
01	.95634	.08303	-.02219	-.11818	-.08321	-.09185
03	.96326	-.04614	-.02242	-.03131	-.11955	-.14277
05	.92992	-.01552	.01896	-.29683	.01861	.11444
08	.93121	-.22389	.21394	.58854	-.03865	-.06637
10	.93913	-.19014	.14838	.01317	.11949	.05164
15	.90093	-.09247	-.11584	-.19669	.27953	.00649
20	.93098	.22173	.09578	-.06054	.05675	-.01592
25	.97025	.66620	-.01831	.31181	.04207	.13699
30	.95743	.08314	.11558	-.07684	-.10308	.12249
35	.93112	-.34796	.01866	-.03446	-.13202	.04062
38	.94669	.15651	.15071	.34358	.05917	.01836
BMB 101	.95698	.09388	.21096	.00165	-.00804	-.07017
103	.86597	-.44896	-.05418	.01387	-.02185	.01428
105	.96812	.04682	.13061	.08565	.07926	-.05553
106	.96522	-.06729	.18835	.00564	.01781	.07578
110	.93596	.05811	.01231	-.07264	.06563	-.21329
112	.95504	.07973	-.03760	.13594	.09911	.10745
113	.92657	-.06276	-.19433	.06387	-.03578	-.02240
115	.94916	-.02795	-.13911	.08255	.02796	-.08309
116	.96822	.14448	.05946	.06760	-.01633	.08813
119	.85735	.08592	-.42746	.06760	-.04112	.07573
120	.93997	-.02309	-.14623	.19090	.02246	-.11501
121	.95316	-.02740	-.09718	-.14475	.04610	-.00662
123	.98192	.04862	-.07023	-.00779	-.01618	.03710
125	.94023	.07411	-.00539	.16844	.05362	.02979
127	.97365	-.05349	.00174	.11033	-.09292	.04576
130	.94490	.07795	.00842	.01514	.07500	-.11218
eigen value	25.5804	.6170	.5192	.3454	.3153	.1998
total load of each factor in %	91.00	2.20	1.85	1.24	1.12	.71

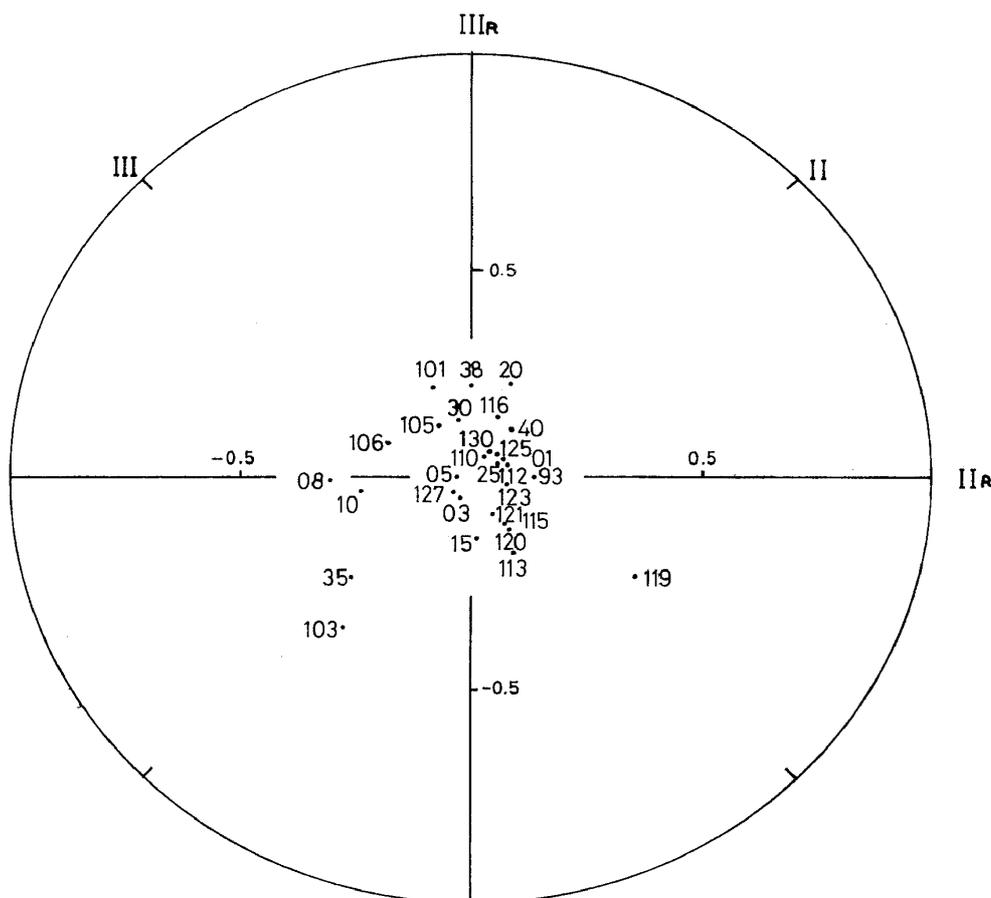


Fig. 11 Benthonic foraminiferal faunas plotted with respect to Factor axes  $II_R$  and  $III_R$ . The numbers are sample numbers.

common component in faunal change of all samples and the second and third factors express 44.4% of the different component in faunal change. In Fig. 11, and faunas of 29 samples are plotted in coordinates projection by mean of the loadings of the second and the third factors. Because the axes of factors by factor analysis are decided by mathematic computation, the direction of the axes do not always fit for interpretation of the faunal changes. To make the interpretation of the faunal change easy, the axes of the second and third factors were rotated  $45^\circ$  dextrally from their original positions II and III to  $II_R$  and  $III_R$  in this projection. In figure 11, the loadings of factor  $II_R$  represent the whole change of the fauna through the stratigraphic horizon and the value differs less when the stratigraphical horizons are closer. The change of the loadings of  $II_R$  with stratigraphic horizon is shown in a column of Fig. 9. From Fig. 9, it is obvious that the value of  $II_R$  is reflected well by the ratio of *C. carinata* to *B. aculeata*. It appears that the majority of the benthonic foraminiferal change can be explained by the ratio of *C. carinata* to *B. aculeata*. According to Aoki (1968), *C. carinata* lived on bottoms shallower than *B. aculeata*. Therefore, the ratio is considered to be a measure to estimate the depth of water in a relative manner.

The ratio *B. aculeata* / *C. carinata* shows the maxima at 260 cm and 50 cm below the key tuff and accordingly  $II_R$  shows the minimal values at these horizons. Hence, it is inferred from the benthonic Foraminifera that water was deepest at these two horizons. The two horizons are close in position to 270 and 65 cm, respectively which represented the highest values of mud contents and only type I sediments, have been judged to be of the

deepest water from the grain-size analysis.

At the horizon in which the ratios show the shallower depths, the sediments tend to have lower mud content and are usually represented by type II and/or III (Fig. 9). This is interpreted to have resulted from the removal of  $4\phi$  grains from the type I, that took place only in shallower depths.

The change in the benthonic foraminiferal fauna occurred in cycles of about 200 cm period and agrees with the change of the sediments. The length of the period also agrees with the period in the change of geomagnetic intensity. At the horizon of the earth's magnetic field reversal no apparent change in benthonic foraminiferal fauna was recognized.

## 2) Faunal change of planktonic Foraminifera

In the present study, only the tests, coarser than 125 microns have been identified. At least 200 planktonic specimens have been identified for each sample, resulting in recording altogether 24 species in 9 genera. In the 29 samples studied, the following species are abundant: *Globigerina pachyderma* (Ehrenberg), *G. falconensis* Blow, *Globigerinoides ruber* (d'Orbigny), *Globorotalia inflata* (d'Orbigny), *Globigerinita glutinata* Egger, *Globigerina immaturus* LeRoy (Table 7).

Factor analysis has been applied to analyse the changes that took place in the planktonic population during the time represented by this section. The analysis is made on the bases of the relative frequencies of 24 species in each of the 29 samples (Table 8). Only the largest three factors, Factor 1, 2, and 3, are considered in the following discussion. It was found that the first factor remained virtually unchanged throughout the section, and the total of its loadings is 93.6% of that of the all factors, indicating that the species composition and their proportion did not change greatly during the time-interval. The total loadings of the second factor is 4.5% of that of the all factors, that of the third factor 0.5%. Since the first factor is virtually constant, the changes observed in the second factor is important, particularly in view of that the second factor expressed the most (70.5%) of the loadings after the first factor has been excluded (Fig. 12). The loadings of the second factor changes greatly in the horizons (Fig. 9). The comparison of the curve with that in the next column (given for the proportion of *G. pachyderma* in the planktonic foraminiferal assemblage) suggests that *G. pachyderma* accounts decisively for the changes. It is revealed, therefore, that the proportion of *G. pachyderma* represents, in effect, the larger part of the difference in the planktonic foraminiferal faunas through the stratigraphic interval.

As shown in Fig. 9, *G. pachyderma* takes about 30% of the planktonic population in their horizons lower than 300 cm below the key tuff; the proportion increases upward to more than 65% at 139 cm after fluctuating at intermediate horizons, e.g. 45% at 285 cm, 39% at 278 cm, all below the key tuff. After attaining 65% at 139 cm below the key tuff, the proportion remains above 60% through the rest of the section up to the key tuff. In this change, the first increase in frequency of *G. pachyderma* at the horizon 285 cm below the key tuff corresponds to the temporal lowering of the paleotemperatures measured for the tests of *G. pachyderma* and *M. barleeanus* (Fig. 9). In the rest of the sequence up to the key tuff, however, the change of paleotemperature does not correspond to the fluctuations shown in the frequencies of *G. pachyderma*.

Fig. 13 shows another result from the factor analysis. It is to determine the species to species relationships that were maintained by planktonic Foraminifera during the accumulation of the present section. In Fig. 13 all planktonic species are plotted in coordinates projection by means of the loadings of the first and second factors. In this projection the point representing *G. pachyderma* is found isolated from the points of the other species, indicating that the frequency of *G. pachyderma* changes in horizon independ-



Table 8 Data of Factor analysis of planktonic foraminiferal faunas.

Sample	FACTOR 1	FACTOR 2	FACTOR 3	FACTOR 4	FACTOR 5	FACTOR 6
KM 93	.90841	.37207	-.01876	.05225	-.00550	-.16754
BM 40	.88437	.44929	-.04197	.08120	-.04743	.00207
01	.79722	.57650	.02179	-.08063	.12122	.07088
03	.85217	.49926	.07199	-.00698	-.12475	.01304
05	.98163	.02629	.01046	.14765	-.02464	.08669
08	.95253	.27293	.00995	-.06551	-.01120	.00615
10	.99074	.04455	.09747	-.01290	.02379	-.07482
15	.90054	-.10861	.04048	-.03608	-.01628	.03003
20	.98275	-.14717	.08014	.03251	.03869	.02775
25	.97856	-.17190	.09930	.01817	-.02088	-.02338
30	.92680	-.12148	.09169	-.00790	.01192	-.02991
35	.98601	-.08466	.11529	-.03779	.04662	-.01363
38	.98125	-.14171	.11769	-.01657	.01754	-.00427
BMB 101	.99156	-.04046	.03998	.00384	.07796	-.05127
103	.98239	-.01579	-.06091	-.12556	-.09152	.03506
105	.98870	-.02395	-.09767	.09089	.00663	-.00010
106	.99056	.07963	-.02267	-.03775	.00296	-.00079
110	.97830	-.16595	-.02330	.10614	-.01575	-.01511
112	.98778	-.07653	-.07944	-.00709	.02979	-.00367
113	.97692	.06783	-.15693	.01326	.10436	.01390
115	.99142	-.01017	-.04206	.00801	.01917	.06379
116	.97993	-.12681	-.12179	-.04045	.01817	-.03659
119	.98497	-.14117	-.02124	-.07850	-.00681	-.03220
120	.98848	-.06168	.06470	.07533	.01714	.05306
121	.98394	-.14945	-.04065	.01415	-.01441	-.02561
123	.98715	-.12038	-.04058	.02646	-.03358	-.00280
125	.99421	-.08750	.01795	-.02359	-.03877	-.00686
127	.97617	-.17579	-.06492	-.08540	-.03779	.00174
130	.98350	-.15714	-.04023	-.01511	-.04641	.03706
eigen value	27.1709	1.2912	.1486	.1061	.0726	.0604
total load of each factor in %	93.63	4.45	.51	.36	.25	.21

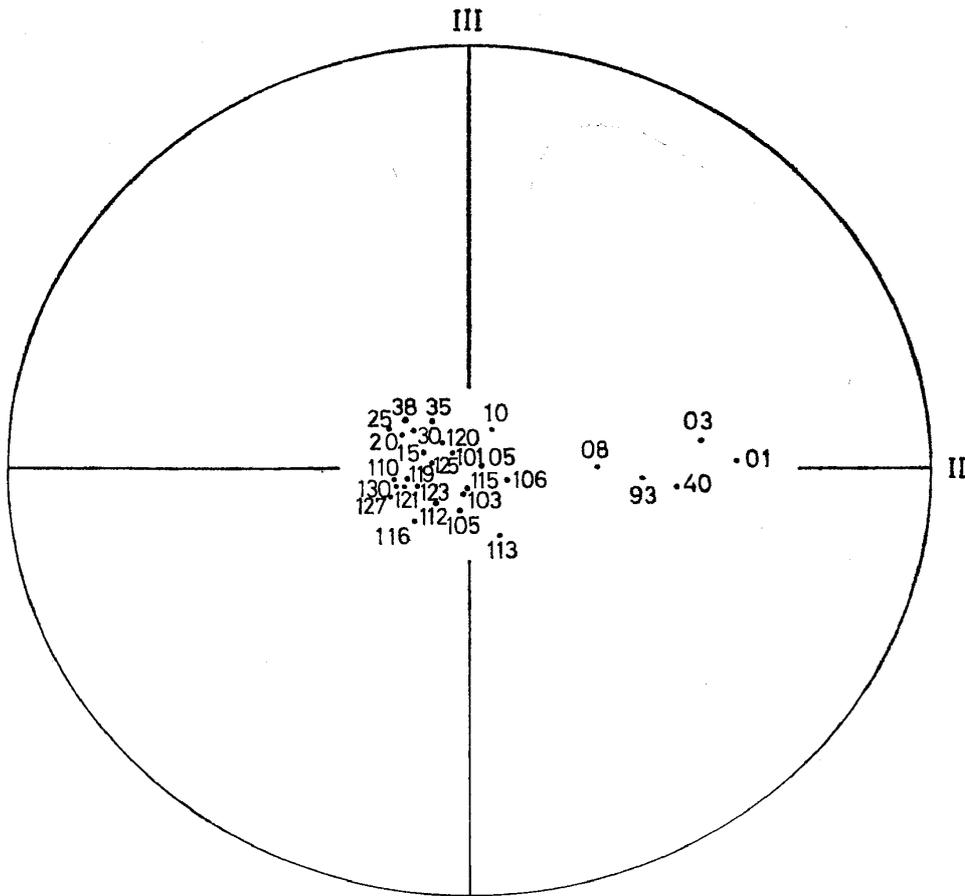


Fig. 12 Planktonic foraminiferal faunas plotted with respect to Factor axes II and III. The numbers are sample numbers.

ently with other species.

Fig. 14 shows the result of plotting on the Motomura's Plane (Niitsuma, 1968), in which a vertical axis is "a" and a horizontal axis is "B" of the values computed by the equation:

$$\log y + a(x-1) = \log B \tag{1}$$

where  $x$  is the rank of frequency of a species in the fauna,  $6$  is the frequency of species in percent,  $a$  and  $B$  are constants calculated by the least square method using the frequencies of six higher ranked species in the fauna. Because the frequency of species,  $y$  is expressed in percentage, the sum of  $y$  of all species in a fauna would be 100. Thus the relation,

$$B = 100 \left( 1 - \frac{1}{10^a} \right) \tag{2}$$

is lead from the equation (1) that gives a curve (Motomura's Line) in Fig. 14. Whether a given fauna fulfills the relation suggested by Motomura (1932) can be determined by the values of  $a$  and  $B$  obtained for the fauna by the equation (1). If a fauna satisfies the relation of logarithmic normal distribution or a fauna is a mixture of faunas which satisfy the relation of equation (1) (Preston, 1948; Nobuhara *et al.*, 1954; Whittaker, 1965), the point for the fauna would be located on the upper leftside of Motomura's Line (Niitsuma, 1968; Niitsuma *et al.*, 1971). Ujiie (1968) reported that the living planktonic Foraminifera in the Indian Ocean satisfied Motomura's relation, though he misused, the term of the

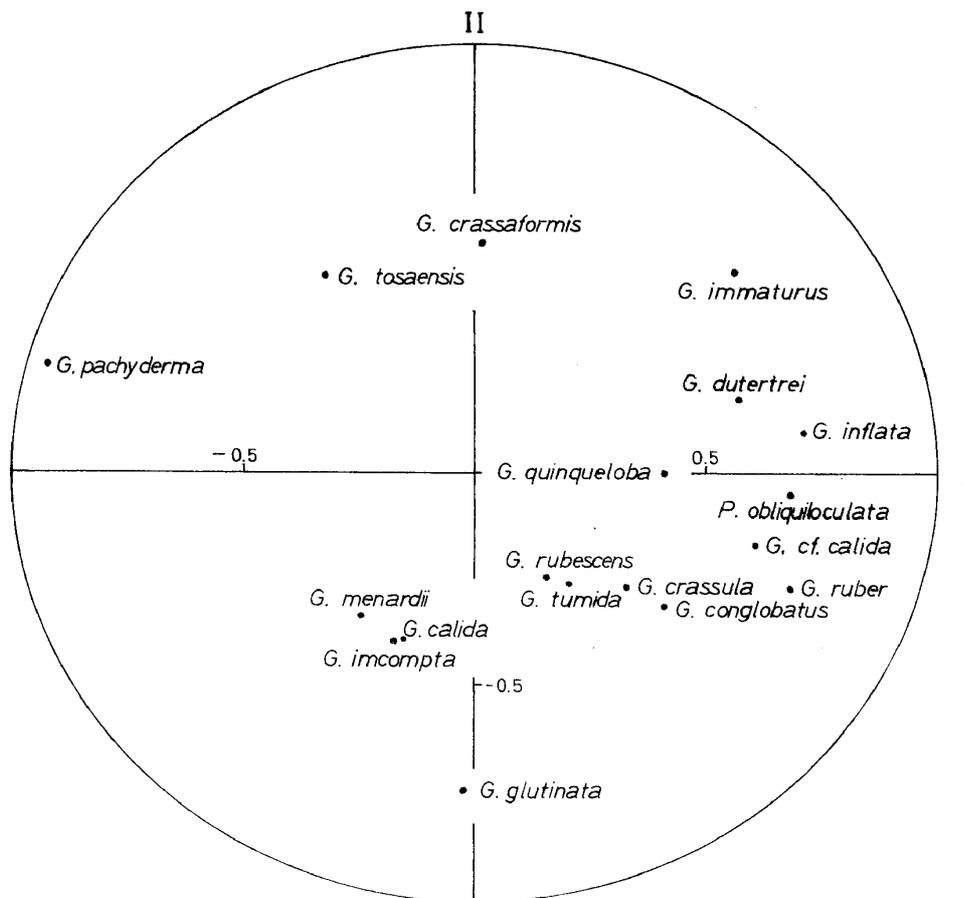


Fig. 13 Planktonic foraminiferal species plotted with respect to Factor axes I and II.

logarithmic normal distribution for this relation. The points obtained for the planktonic foraminiferal faunas treated in this study, however, do not fall near Motomura's Line but scatter widely in the upper region (Fig. 14). When the plots are made excluding *G. pachyderma* from the faunas points are well aligned along Motomura's Line (Fig. 15). However the results are interpreted as suggesting that *G. pachyderma* belonged to a different ecosystem from that of other planktonic foraminiferal species. This interpretation is born out by the other data: 1) the proportion of *G. pachyderma* changes in horizon independently with other species (Fig. 13); 2) paleotemperatures derived from the tests of *G. pachyderma* differ largely and show different change from that of *P. obliquiloculata* (Fig. 9). The interpretation seems to be quite reasonable also in view of the modern geographical distribution of these two species: according to Bradshaw (1959), *G. pachyderma* is a representative of the cold water mass; *P. obliquiloculata* is of the warm water mass to which group the other species occurring in the present section belong.

The mixture of cold and warm water species in a thanatocoenosis from every rock sample can be explained by the migration at the time of the boundary between cold and warm currents in the depositional environment. It is maintained here that the composition of the thanatocoenosis in the present section depended upon the duration of the current that dominated the area, rather than the fluctuation of faunal composition through time that took place within a single water mass. It appeared that a slight change in water temperature took place in *G. pachyderma* carrying cold water at 285 cm below the key tuff, and after that the temperature remained virtually constant. The general trend of increase of the

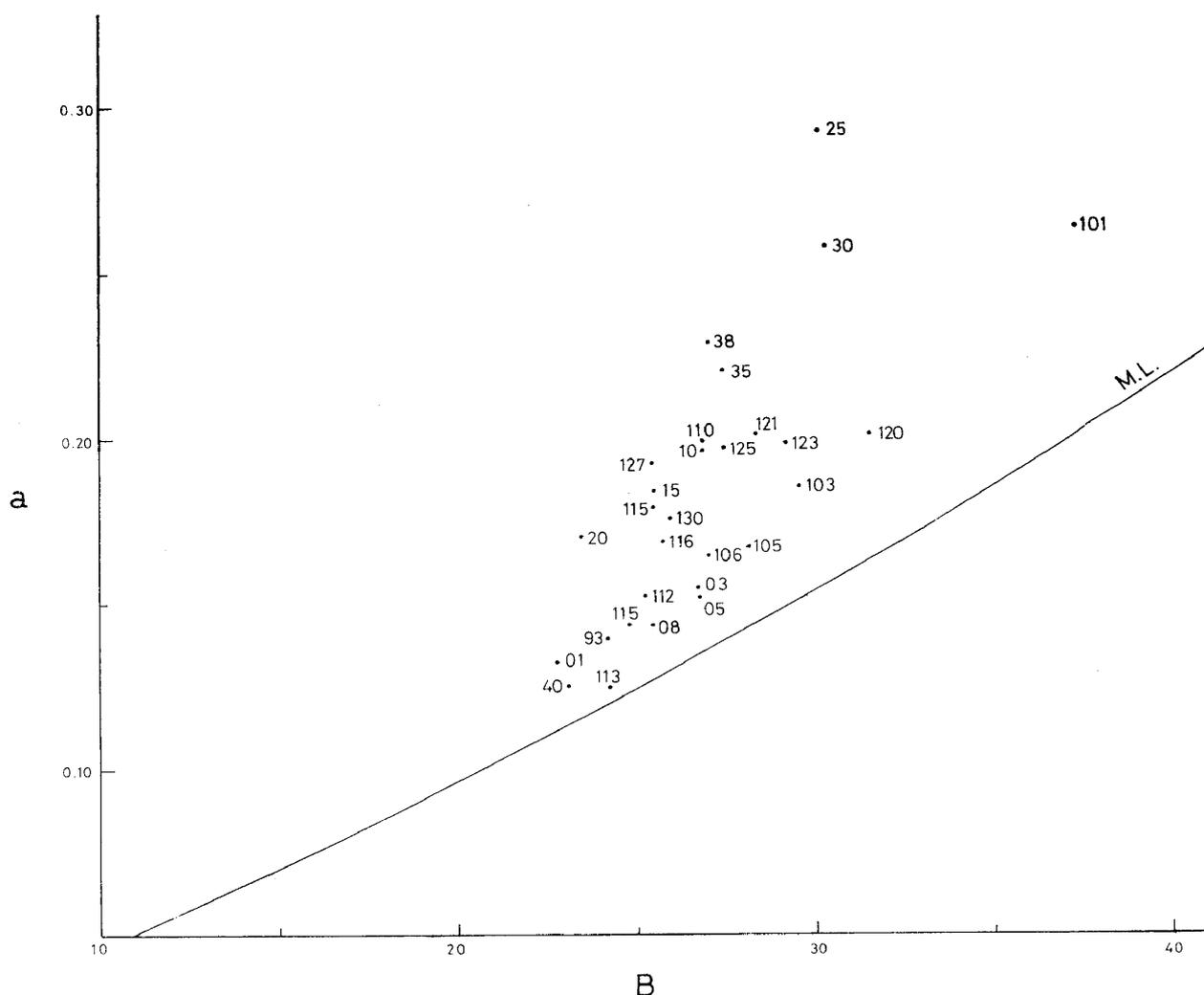


Fig. 14 Plots of  $a$  and  $B$  obtained for 29 planktonic foraminiferal faunas on the "Motomura's Plane".

proportion of *G. pachyderma* commenced from 255 cm level is the reflection of the change that cold water current occupied the area for a longer time than before. The increasing frequency of *G. pachyderma* is found between 255 and 139 cm below the key tuff and this period is coincident with that of the geomagnetic reversal between 240 and 100 cm below the key tuff.

#### CONCLUSION

The behavior of the earth's magnetic field during reversing at the boundary between the Matuyama and Brunhes Polarity epochs has been revealed in detail by measuring DRM of the sedimentary sequence accumulated with high rate of sedimentation. The results show that the time of duration in which the field completed change of its polarity was approximately 4700 years. Intensity of the magnetic field fluctuated with a period of approximately 6700 years around the reversal and attained its minimum value at a reversing time. During the polarity change the position of the virtual magnetic pole drifted along the meridian of  $120^{\circ}$  E.

The depth of water under which the sediments accumulated appears to have remained within 200–300 m; the depth estimate is based on the finding that the

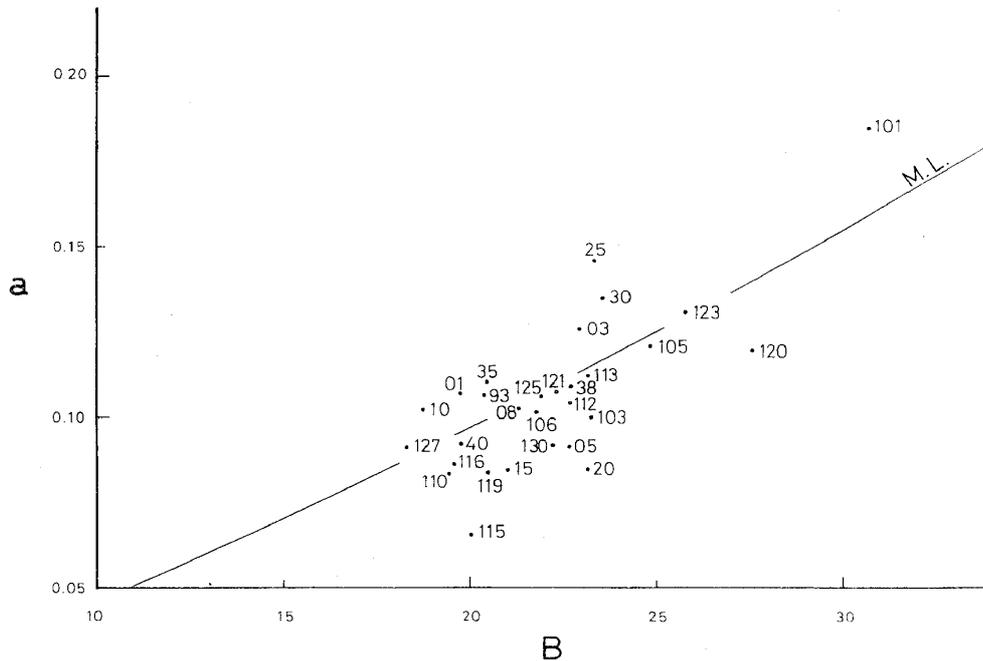


Fig. 15 Plots of  $a$  and  $B$  obtained for 29 planktonic foraminiferal faunas without *Globigerina pachyderma* on "Motomura's Plane".

paleotemperature differed by  $10^{\circ}\text{C}$  between the surface and the bottom of the water-column during the deposition. Within this rather narrow range, the depth of water fluctuated with a period of approximately 7000 years; its phase, however, does not appear to have been related particularly to the magnetic field reversal. The change in water temperature in *P. obliquiloculata* carried by a warm current is judged not to have been appreciable, on the basis of the isotopic measurements on the tests of this species. No appreciable change has been detected for the cold water current which also occupied the area during the accumulation of the sedimentary sequence, except a decrease by  $3^{\circ}\text{C}$  detected at one horizon from the isotopic measurements of the tests of a planktonic species *G. pachyderma*. This horizon of a temporal cooling has also been detected by the benthonic species *M. barleeanus*, also measured isotopically, and is estimated to have taken place approximately 1000 years before the field reversal commenced. Whether this cooling was related causally to the magnetic reversal is still open to question.

Change simultaneous with the magnetic field reversal has been found in the relative frequencies of planktonic Foraminifera. It is particularly represented by a cold water planktonic species, *G. pachyderma*, whose proportion increased gradually from 30% before the reversal to as much as 60% after the completion of the reversing, while the temperature of water remained virtually unaltered. The increase in proportion is, therefore, interpreted to reflect a longer stay of the cold water-masses in the locality during and after the reversal than before.

Finally various lines of evidence indicate that the magnetic field reversal at the boundary of Matuyama-Brunhes Polarity epochs can be related to the change in the circulation pattern of superficial water-masses that caused an influx of cold water into this area in the northwestern Pacific compared with the period before the geomagnetic polarity change.

Table 9 Data of Factor analysis of planktonic species.

species	FACTOR 1	FACTOR 2	FACTOR 3	FACTOR 4	FACTOR 5	FACTOR 6
Globigerina bulloides	.21047	.47648	-.28544	-.03583	-.10679	.25449
G. calida	-.15186	-.39621	.58987	-.13708	-.27415	.25846
G. falconensis	.85660	.14589	.04183	-.12626	.22136	.01177
G. pachyderma	-.91940	.25399	.00593	.12693	.03561	-.08642
G. quinqueloba	.41364	.00164	-.52674	.22709	-.21839	.13829
G. rubescens	.15911	-.24882	.09901	.11059	.52907	.03038
Globigerinita glutinata	-.02955	-.75174	-.11247	-.15785	.17118	.23278
G. uvula	.02419	-.17895	.01328	-.04991	-.57442	-.50319
Globigerinoides conglobatus	.41372	-.31057	.06967	.53441	.05866	.28851
G. immaturus	.56217	.47061	-.29919	.11647	-.25624	.34734
G. ruber	.68362	-.27081	.08301	.24286	-.20165	-.02618
Globoquadrina dutertriae	.57149	.17600	.14882	.10402	-.53345	-.00579
Globorotalia crassaformis	.01845	.54253	.30225	-.35069	-.16952	.45445
G. crassula	.32848	-.26737	-.30882	-.00847	.09794	-.28018
G. inflata	.71287	.09338	.02738	-.10452	.22951	-.17166
G. menardii	-.24361	-.35575	-.41730	-.24169	.26586	.18421
G. tosaensis	-.32519	.46475	.44414	.16358	.22338	.06468
G. tumida	.20555	-.26072	.66155	-.29884	-.07306	.04176
Pulleniatina obliquiloculata	.68510	-.05003	-.10732	-.40314	.19351	.03429
Orbulina universa	-.30065	-.26224	-.29546	-.38117	-.30956	.05057
Sphaeroidinella dehisceus	-.31805	-.28187	.02752	-.39807	-.11802	.50323
Globigerina cf. calida	.61015	-.16618	.46436	-.21291	.06341	-.17505
Globigerinella siphonifera	.15255	-.34132	-.02039	.68919	-.06737	.44245
Globigerina incompta	-.14750	-.39495	-.07062	.16910	-.47631	-.08456
eigen value	5.0557	2.7870	2.1718	1.8359	1.8176	1.5161
total load of each factor in %	20.95	11.62	9.04	7.64	7.56	6.51

## APPENDIX

PROGRAM OF ELECTRONIC COMPUTER FOR DATA OF  
PALEOMAGNETIC MEASUREMENT

Numerical analyses of data obtained from the present paleomagnetic measurements were processed through a high speed numerical calculation by the Electronic Computer, NEAC 2200-Model 500 at the Computer Center of Tohoku University. In this Appendix, the processes of the analysis and program for electronic computer are described.

A series of values had been obtained by measurement at every  $30^\circ$  around three orthogonal spin axes of one sample. From them the error caused by linear drift of the magnetometer was corrected by using the values of the first and last measurement,  $A_1^m$  and  $A_{13}^m$ , in a spin for the fixed axis.

$$Ai = \frac{(A_1^m - A_{13}^m)(i-1)}{12} + Ai^m$$

where,  $Ai$  is the value after this correction, and  $Ai^m$  is  $i$ -th measured value in a spin.

The measured value is generally distributed along a sine curve when it was plotted. Two orthogonal components,  $Va$  and  $Vb$ , of the projection of the magnetic moment upon a plane normal to the spin axis are calculated through Fourier analysis:

$$Va = \frac{2}{12} \sum_{i=1}^{12} Ai \cos \{30^\circ \times (i-1)\}$$

$$Vb = \frac{2}{12} \sum_{i=1}^{12} Ai \sin \{30^\circ \times (i-1)\}$$

The standard deviation,  $SD$ , is determined as a ratio of mean distance of the plotted intensities  $Ai$  versus a sine curve calculated from measured values:

$$SD = \frac{\sqrt{\frac{1}{12} [Va \cos \{30^\circ(i-1)\} + Vb \sin \{30^\circ(i-1)\} - Ai + \overline{Ai}]^2}}{\sqrt{Va^2 + Vb^2}}$$

where,  $\overline{Ai}$  is average value of the measured value of a spin,  $A_1 \dots A_{12}$ , and  $\sqrt{Va^2 + Vb^2}$  is an amplitude of the sine curve.

Three pairs of orthogonal components of magnetic moment and three standard deviations of each spin are calculated, and from those values the weighted average of three orthogonal components of magnetic moment by the standard deviation are calculated by the equation:

$$V = \frac{SD_2 V_1 + SD_1 V_2}{SD_1 + SD_2}$$

where,  $V$  is the averaged value;  $SD_1$  and  $SD_2$  are the standard deviations, respectively of measurements from which  $V_1$  and  $V_2$  are calculated.

The three orthogonal components of magnetic moment of a sample are translated into the components,  $V_{NS}$ ,  $V_{EW}$  and  $V_{UD}$ , at the present sampling position and these at the time when the sediments had accumulated, by rotation of the coordinated, taking into the account of the directions of the sample and of the bedding plane. Declination  $D$ , and inclination  $I$ , of magnetic moment are calculated from the components:

$$\tan I = \frac{V_{UD}}{\sqrt{V_{EW}^2 + V_{NS}^2}}$$

$$\tan D = \frac{V_{EW}}{V_{NS}}$$

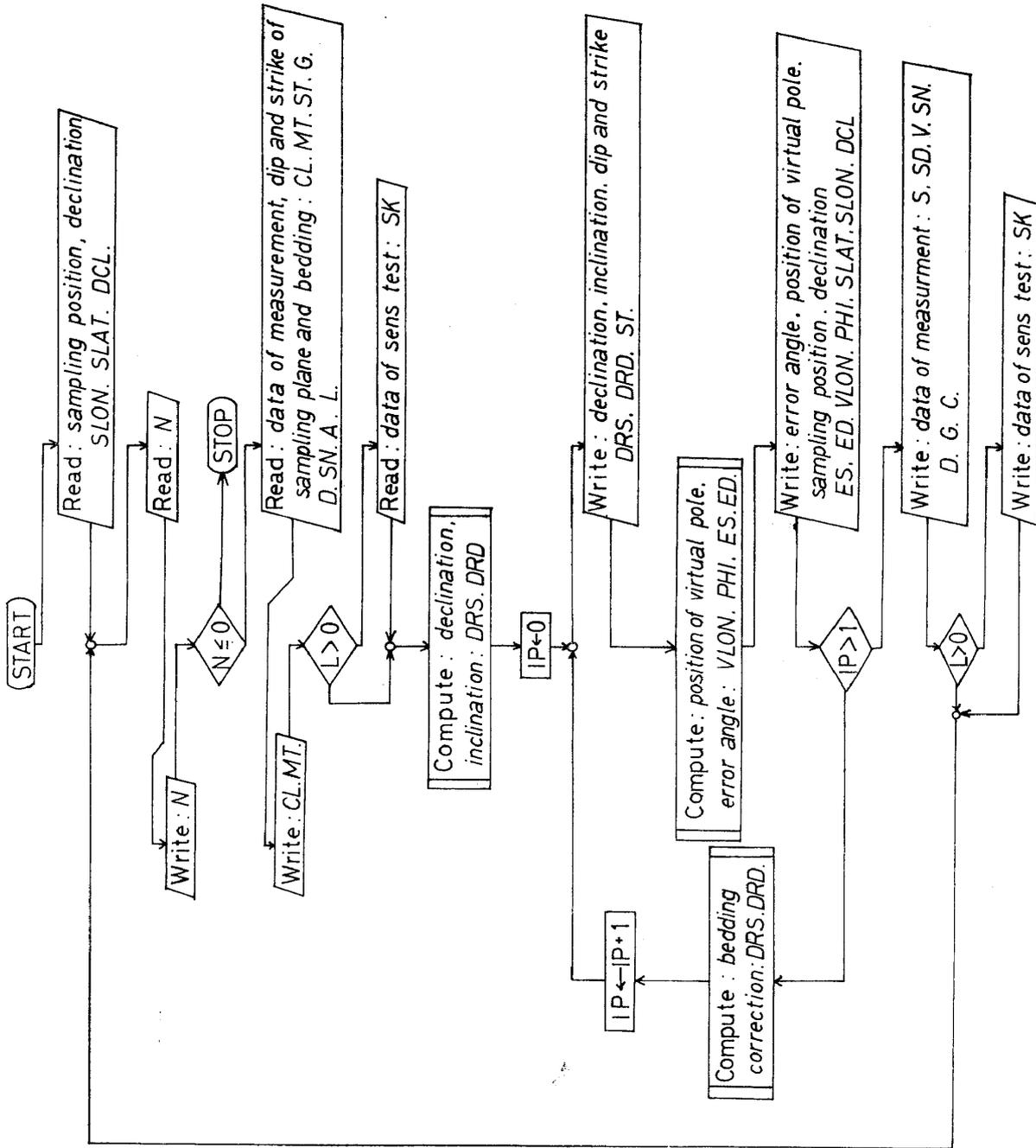


Fig. 16 Flow chart of the program for analysis of paleomagnetic data.

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C      ANALYSIS OF PALEOMAGNETIC DATA
      DIMENSION D(3),ST(2,2),MT(2),SN(3),A(13,3),SK(14),V(5,3)
      DIMENSION VA(3),VB(3),SD(5),C(13,3)
      READ(11,87) SLON,SLAT,DCL
87  FORMAT(3F5.1)
      1 READ(11,80) N
80  FORMAT(I5)
      IF(N) 2,2,3
      3 READ(11,81) CL,(MT(I),I#1,2),((ST(I,J),I#1,2),J#1,2),G
81  FORMAT(A8,I12,I10,4F5.1,F6.1)
      READ(11,82) (D(J),J#1,3),(SN(J),J#1,3)
82  FORMAT(3F5.1,3E10.2)
      READ(11,83) ((A(I,J),I#1,13),J#1,3)
83  FORMAT(9F7.1)
      READ(11,84) L
84  FORMAT(I4)
      IF(L) 35,35,34
34  READ(11,85) SK(1)
85  FORMAT(E10.1)
      READ(11,86) (SK(I),I#2,14)
86  FORMAT(8F10.1)
35  CONTINUE
      WRITE(15,51) N
      WRITE(15,52) CL,(MT(I),I#1,2)
      DO 5 J#1,3
      C(13,J)#A(13,J)
      DO 6 I#1,12
      C(I,J)#A(I,J)
      A(I,J)#((A(1,J)-A(13,J))*FLOAT(I-1))/12.0+A(I,J)
6  CONTINUE
      VA(J)#((A(1,J)-A(7,J))+0.8666*(A(2,J)-A(6,J)-A(8,J)+
1A(12,J))+0.5*(A(2,J)-A(5,J)-A(9,J)+A(11,J)))*SN(J)*
2D(J)**3.0)/6.0/G
      VB(J)#((A(4,J)-A(10,J))+0.8666*(A(3,J)+A(5,J)-A(9,J)-
1A(11,J))+0.5*(A(2,J)+A(6,J)-A(8,J)-A(12,J)))*SN(J)*
2D(J)**3.0)/6.0/G
      SA#0.0
      SD(J)#0.0
      DO 40 I#1,12
      A(I,J)#A(I,J)/G*SN(J)*D(J)**3.0
      SA#SA+A(I,J)/12.0
40  CONTINUE
      DO 12 I#1,12
      SD(J)#SD(J)+(ABS(SA+VA(J))*COS(30.0/57.296*FLOAT(I-1))+
1VB(J))*SIN(30.0/57.296*FLOAT(I-1))-A(I,J))**2.0/12.0
12  CONTINUE
      SD(J)#SQRT(SD(J))/SQRT((ABS(VA(J))**2.0+(ABS(VB(J)))
1**2.0)
5  CONTINUE
      V(2,1)#VB(3)
      V(2,2)#VB(1)
      V(2,3)#VB(2)
      V(3,1)#VA(1)
      V(3,2)#VA(2)
      V(3,3)#VA(3)

```

Fig. 17a

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DO 13 J#1,3
SD(4)#SD(1)
SD(5)#SD(2)
JPL2#J+2
V(1,J)#V(2,J)*SD(J)/(SD(J)+SD(JPL2))+V(3,J)*SD(JPL2)/
1(SD(J)+SD(JPL2))
V(4,J)#(V(2,J)+V(3,J))/2.0
V(5,J)#ABS(V(2,J)-V(3,J))/2.0
13 CONTINUE
IP#0
DPS#ABS(ST(2,1))
DO 14 I#1,4
VTEMP#V(I,2)
V(I,2)#VTEMP*SIN(DPS/57.296)+V(I,3)*COS(DPS/57.296)
V(I,3)#-VTEMP*COS(DPS/57.296)+V(I,3)*SIN(DPS/57.296)
14 CONTINUE
STS#ST(1,1)/57.296
DO 18 I#1,4
VTEMP#V(I,1)
V(I,1)#ST(2,1)/ABS(ST(2,1))*(-VTEMP*COS(STS)-V(I,2)*
1SIN(STS))*ST(1,1)/ABS(ST(1,1))
V(I,2)#ST(2,1)/ABS(ST(2,1))*(VTEMP*SIN(STS)-V(I,2)*
1COS(STS))*ST(1,1)/ABS(ST(1,1))
18 CONTINUE
19 DRD#ATAN(V(1,3)/SQRT((ABS(V(1,1)))**2.0+(ABS(V(1,2)))
1**2.0))
70 IF(DRD) 71,72,73
71 P#-ATAN(2.0*COS(DRD)/SIN(DRD))
GO TO 74
72 DRD#DRD+0.0001
GO TO 70
73 P#3.1416-ATAN(2.0*COS(DRD)/SIN(DRD))
74 DRD#DRD*57.296
SLAT#SLAT*57.296
ISN#0
26 IF(V(1,1)) 22,23,24
22 DRS#-ATAN(V(1,2)/V(1,1))+DCL/57.296
IF(ABS(DRS)-1.5708) 41,41,42
41 SPHI#SIN(SLAT)*COS(P)+COS(SLAT)*SIN(P)*COS(DRS)
ISN#ISN+1
75 PHI#ATAN(SPHI/SQRT(1.0-(ABS(SPHI))**2.0))
SDLON#SIN(P)*SIN(DRS)/COS(PHI)
DLON#ATAN(SDLON/SQRT(1.0-(ABS(SDLON))**2.0))
CPC#SIN(SLAT)*SIN(PHI)+COS(SLAT)*COS(PHI)*COS(DLON)
COSP#COS(P)
IF(ABS(CPC-COSP).LE.0.01) GO TO 45
DLON#(DLON/ABS(DLON))*(3.1416-ABS(DLON))
45 VLON#SLON-DLON*57.296
DRS#DRS*57.296
PHI#PHI*57.296
SLAT#SLAT*57.296
IF(1-ISN) 76,76,77
76 WRITE(15,53)DRS,DRD,ST(1,1),ST(2,1)
GO TO 25
42 DRS#(DRS/ABS(DRS))*(3.1416-ABS(DRS))

```

Fig. 17b

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GO TO 43
23 V(1,1)#V(1,1)+0.0001
GO TO 26
24 DRS#ATAN(V(1,2)/V(1,1))-DCL/57.296
IF(1.5708-ABS(DRS)) 44,43,43
43 SPHI#SIN(SLAT)*COS(P)-COS(SLAT)*SIN(P)*COS(DRS)
GO TO 75
44 DRS#(DRS/ABS(DRS))*(3.1416-ABS(DRS))
GO TO 41
77 WRITE(15,54) DRS,DRD,ST(1,1),ST(2,1)
25 DO 31 J#1,3
V(5,J)#ABS(V(2,J)-V(3,J))/2.0
31 CONTINUE
IF(V(4,1)) 38,39,38
39 V(4,1)#V(4,1)+0.0001
38 ES#ABS(ATAN(V(4,2)/V(4,1))-ATAN((V(4,2)+V(5,2))/
1(V(4,1)-V(5,1))))*57.296
ED#ABS(DRD-ATAN((V(4,3)+V(5,3))/SQRT((ABS(V(4,1))-
1V(5,1)))*2.0+(ABS(V(4,2)-V(5,2)))*2.0))*57.296)
WRITE(15,55) ES,ED
WRITE(15,64) PHI,VLON,SLAT,SLON,DCL
IF(IP-1) 28,29,29
28 DO 27 I#1,4
STS#ST(1,2)/57.296
ASTS#ABS(STS)
VTEMP#V(I,1)
V(I,1)#VTEMP*SIN(ASTS)+V(I,2)*COS(ASTS)*STS/ASTS
V(I,2)#V(I,2)*SIN(ASTS)-VTEMP*COS(ASTS)*STS/ASTS
DPS#ST(2,2)/57.296
VTEMP#V(I,1)
V(I,1)#VTEMP*COS(DPS)-V(I,3)*SIN(DPS)
V(I,3)#VTEMP*SIN(DPS)+V(I,3)*COS(DPS)
VTEMP#V(I,1)
V(I,1)#-V(I,2)*COS(ASTS)*ASTS/STS+VTEMP*SIN(ASTS)
V(I,2)#VTEMP*COS(ASTS)*ASTS/STS+V(I,2)*SIN(ASTS)
27 CONTINUE
ST(1,1)#ST(1,2)
ST(2,1)#ST(2,2)
IP#IP+1
GO TO 19
29 S#SQRT((ABS(V(1,1)))*2.0+(ABS(V(1,2)))*2.0+
1(ABS(V(1,3)))*2.0)
WRITE(15,56) S
WRITE(15,57)(SD(J),J#1,3)
WRITE(15,58)((V(J,I),I#1,3),J#1,3)
WRITE(15,59)(SN(I),I#1,3)
WRITE(15,60)(D(I),I#1,3),G
WRITE(15,61)((C(I,J),I#1,13),J#1,3)
IF(L) 36,36,37
37 WRITE(15,62) SK(1)
WRITE(15,63) (SK(I),I#2,14)
51 FORMAT(1H1,5X,7HJOB NO.,2X,15)
52 FORMAT(1H0,10HSAMPLE NO.,2X,A8/1H0,5HSTART,2X,110,1X,
12HTO,1X,16)

```

Fig. 17c

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53 FORMAT(1H0,42HNORTH POLE DIRECTION STRIKE DIP (W+,E-) N,
12F7.2,/1H,42HSTRIKE DIP OF SAMPLE OR BED (W+,E-) N,
22F7.2)
54 FORMAT(1H0,42HNORTH POLE DIRECTION STRIKE DIP (W+,E-) S,
12F7.2,/1H,42HSTRIKE DIP OF SAMPLE OR BED (W+,E-) S,
22F7.2)
55 FORMAT(1H,42HERROR ANGLE OF STRIKE AND DIP (+-DEGREE) ,
12F7.2)
56 FORMAT(1H0,32HMAGNETIC INTENCITY (OE*CM**3/CC),10X,
11PE11.4)
57 FORMAT(1H0,32HSTANDARD DEVIATION (N-E,E-D,D-N)/1H,25X,
13F12.5)
58 FORMAT(1H0,44HLAST DIRECTION VECTOR (OE*CM**3/CC,SN WE UD)
1,/1H,25X,3(3(1PE12.4)/1H))
59 FORMAT(1H0,40HSENS OF MAGNETMETER (OE/MM,N-E,E-D,D-N),/
11H,25X,3(1PE12.4))
60 FORMAT(1H0,25HDISTANS (CM,N-E E-D D-N),3F12.2/1H0,
121HVOLUME OF SAMPLE (CC),6X,F10.2)
61 FORMAT(1H0/1H0,23HDATA OF MESUREMENT (MM),/1H,3HN-E,/
11H,9F8.2/1H,4F8.2/1H,3HE-D,/1H,9F8.2/1H,4F8.2/1H,
23HD-N,/1H,9F8.2/1H,4F8.2)
62 FORMAT(1H0/1H0,25HDATA OF SENS TEST (OE,MM),/1H,1PE11.4)
63 FORMAT(1H,9F8.2/1H,4F8.2)
64 FORMAT(1H0,42HVIRTUAL MAGNETIC POLE POSITION LAT, LON(NE),
12F7.2/1H,38HSAMPLING POSITION AND DECLINATION(W+),
23F7.2)
36 GO TO 1
2 STOP
END

```

Fig. 17d

Fig. 17 Program of electronic computer for analysis of paleomagnetic data.

where,  $V_{NS}$ ,  $V_{EW}$  and  $V_{UD}$  are components of north-south, east-west and up-down directions, respectively.

Error angles of declination  $ED$ , and of inclination  $EI$ , are calculated from the differences of two components in the same direction:

$$\tan(I-EI) = \frac{V_{UD} + \Delta V_{UD}}{\sqrt{(V_{EW} - \Delta V_{EW})^2 + (V_{NS} - \Delta V_{NS})^2}}$$

$$\tan(D-ED) = \frac{V_{EW} + \Delta V_{ED}}{V_{NS} + \Delta V_{NS}}$$

where,  $V$  is the average value of components of the same direction and  $\Delta V$  is the difference of the components of the same direction.

The virtual magnetic north pole position ( $\phi'$ ,  $\lambda'$ ) is calculated from declination  $D$ , inclination  $I$ , latitude and longitude of sampling position ( $\phi_0$ ,  $\lambda_0$ ):

$$\cot P = (1/2) \tan I$$

$$\sin \phi' = \sin \phi_0 \cos P + \cos \phi_0 \sin P \cos D$$

$$\sin(\lambda_0 - \lambda') = \sin P \sin D / \cos \phi'$$

Flow chart of this calculations and program of electronic computer are shown in Figs. 16 and 17.

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